

MULTIPLER BASED TECHNIQUE TO DERIVE ATMOSPHERIC WINDS FROM KALPANA-1

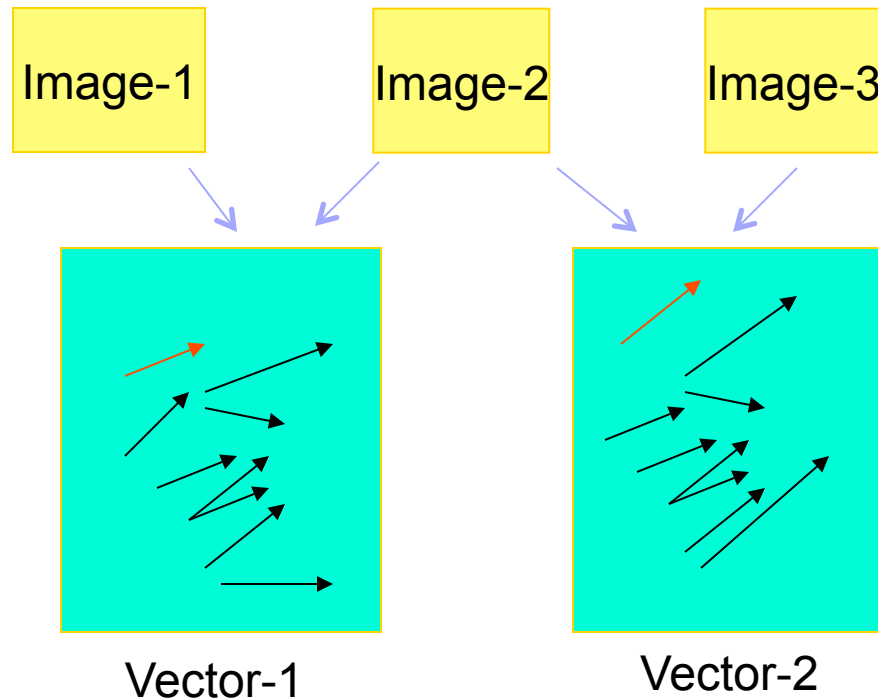
Sanjib K. Deb, C. M. Kishtawal, Inderpreet Kaur,
P. K. Pal and A. S. Kirankumar

Space Applications Centre
Indian Space Research Organization,
Ahmedabad, India
E-mail: sanjib_deb@sac.isro.gov.in

Contents:

- Limitations of Triplet algorithm.
- Temporal scale of AMV.
- Recent improvement in retrieval algorithm.
- Validation statistics with radiosonde.
- Spatial error analysis with NCEP GFS.
- Findings of this study.
- Future directions.

Conventional Triplet Method : Limitation



- Conventional methodology requires that a vector is available in both sets, if not so, such vectors are rejected, because they don't get "support".
- Many "isolated" vectors thus get eliminated, even though they represent the real situation.

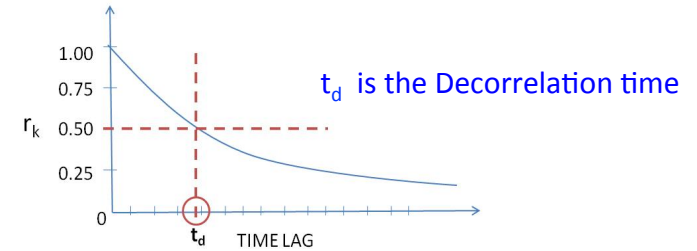
❖ For a discrete time series $\{u_i : 1 \leq i \leq n\}$, for every member of the series some degree of dependence upon the preceding members can be deduced.

❖ For time lag k , autocovariance :
$$c_k = \frac{1}{n-k-1} \sum_{i=k+1}^n (u_i - \bar{u})(u_{i-k} - \bar{u})$$

❖ **Autocorrelation function** : $r_k = \frac{c_k}{c_0}$, c_0 is the variance of the time series.

❖ Defining **Vector wind** is a complex quantity $U = u+iv$, where i is square root of -1

❖ Vector Autocorrelation Coefficient :
$$c_k = \frac{\frac{1}{n} \sum_{i=1}^{n-k} (X_i - \bar{X})(Y_i - \bar{Y})^*}{\left[\frac{1}{n} \sum_{i=1}^{n-k} |X_i - \bar{X}|^2 \right]^{\frac{1}{2}} \left[\frac{1}{n} \sum_{i=1}^{n-k} |Y_i - \bar{Y}|^2 \right]^{\frac{1}{2}}}$$
 , where X and Y are two complex quantities.



❖ **DECORRELATION TIME** is an Index to define the time at which the Autocorrelation coefficient drops to 0.5 .

❖ Assumed beyond this time the two variables cease to have any meaningful relationship.

REGION OF STUDY:

❖ Indian Monsoon Region

❖ To ensure availability of reasonable temporal observations those regions were selected where data gaps were less than a certain threshold.

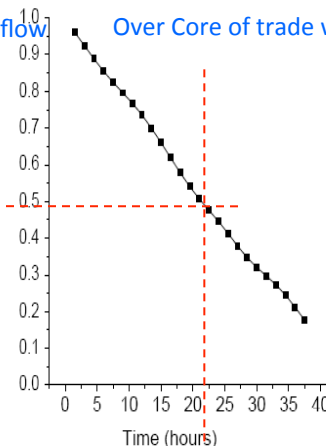
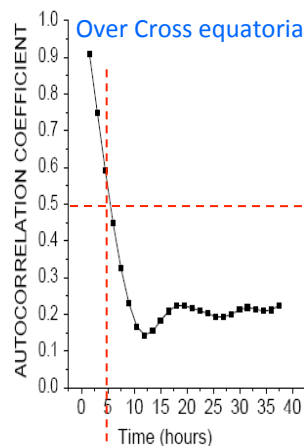
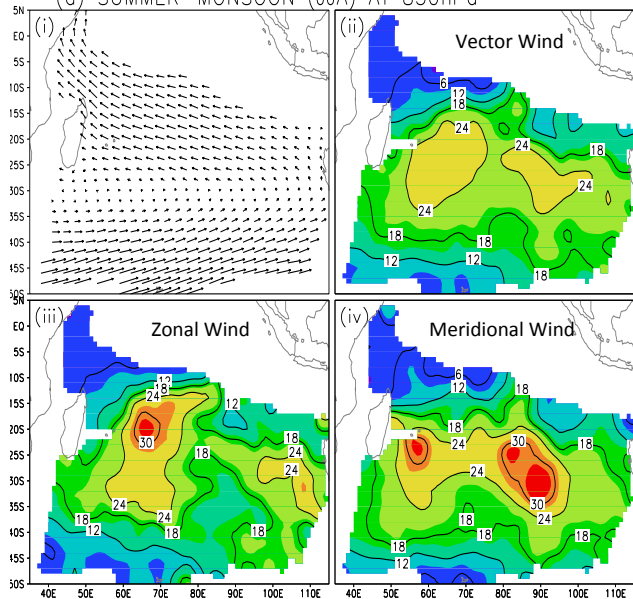
❖ Threshold for upper atmosphere was fixed at 50% of maximum available observations.

❖ At lower level it was taken to be 20- 25 % of maximum available observations.

Decorrelation Time scales for Summer Monsoon (JJA)

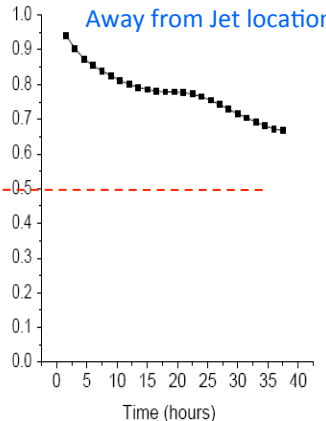
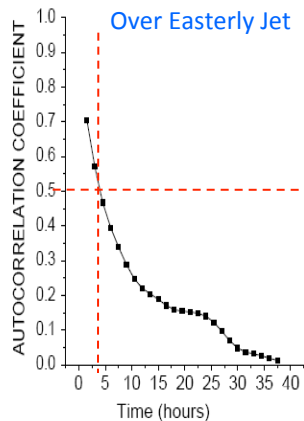
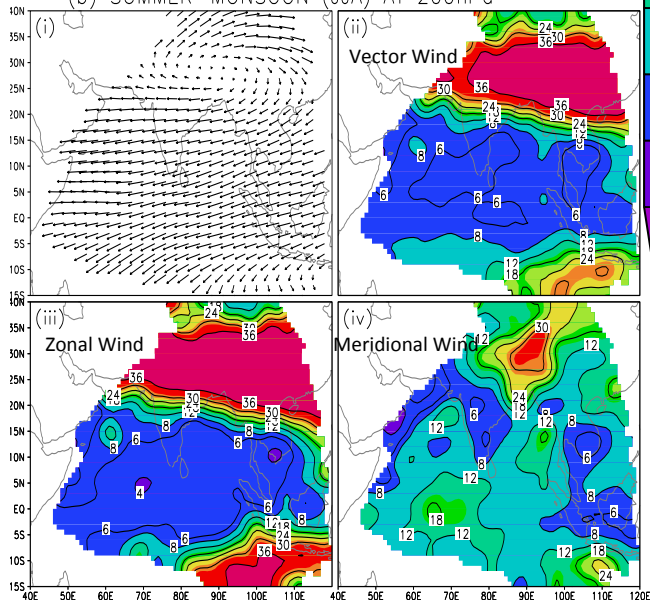


(a) SUMMER-MONSOON (JJA) AT 850hPa



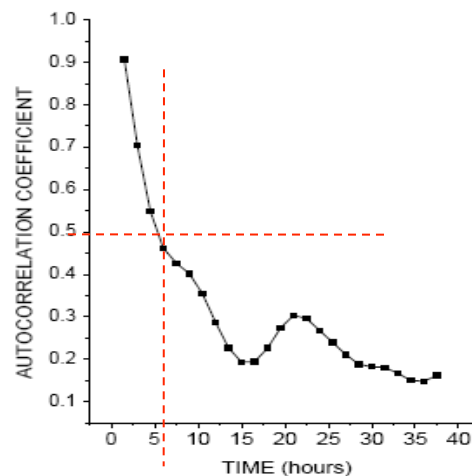
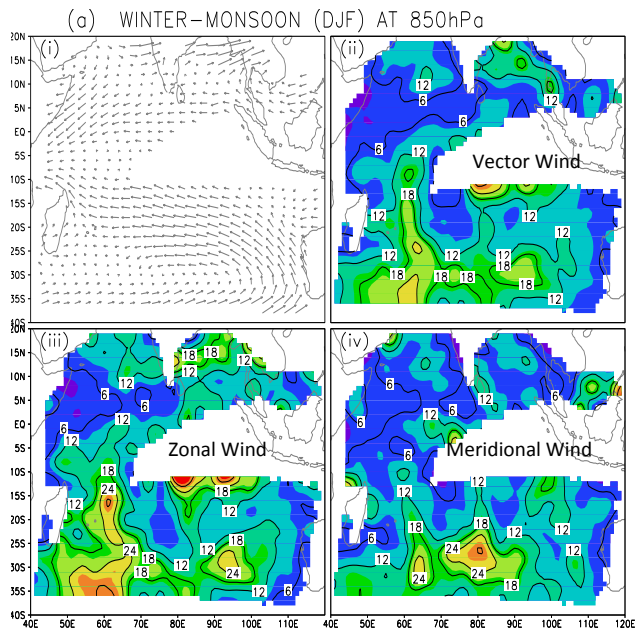
- Most stable winds occur in core of the trades for 850hPa winds.
- This persistence of winds is evident in $t_d \sim 18-24$ hours.
- Weakest winds occur in region over northern tip of Madagascar.
- In this region $t_d \sim 6$ hours.

(b) SUMMER-MONSOON (JJA) AT 200hPa



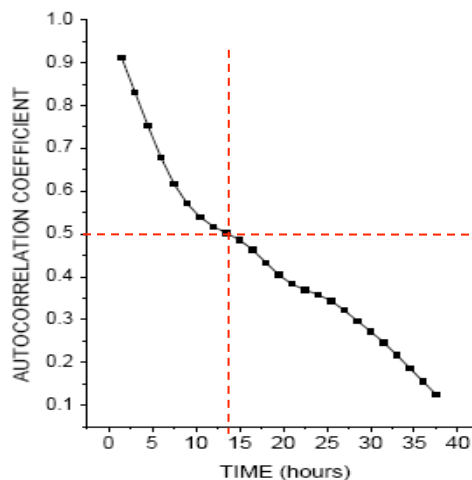
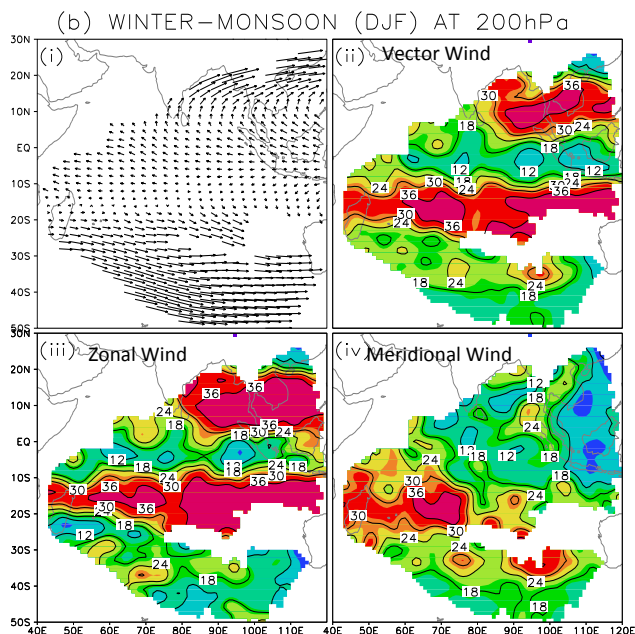
- Winds over Northern IO and peninsular India have low $t_d \sim 4-8$ hours.
- Strong lateral and vertical wind shears in the Easterly jet ($\sim 14N$) might be responsible for low stability.
- Highly stable zonal and meridional flow over the mid-latitudes, $t_d \sim 24-36$ hours.

Decorrelation Time scales for Winter Monsoon (DJF)



▪ Cloud free conditions lead to less data over the subcontinent.

▪ Persistence levels of the order of 6-12 hours were observed in all the three cases.



▪ Westerlies towards the Australian Region are stable and strong winds.

▪ Very high decorrelation time scales ~36 hours were observed

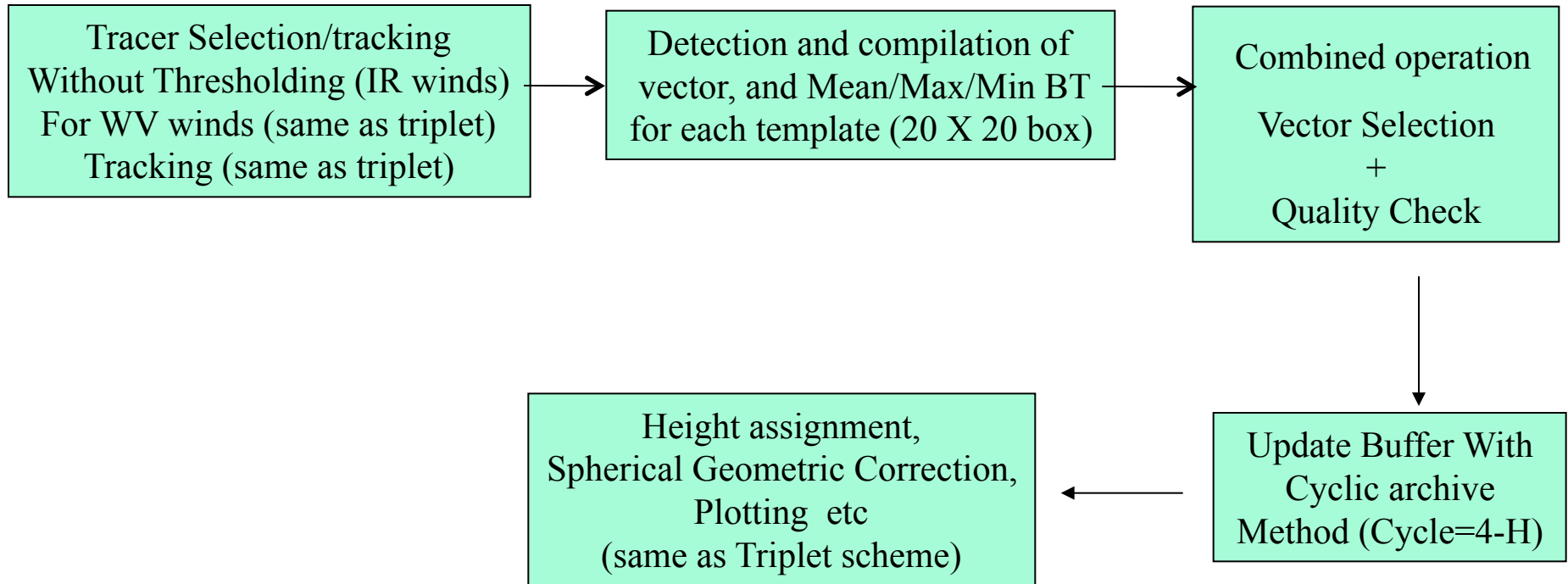
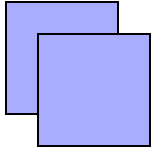
Findings from temporal scale satellite winds study



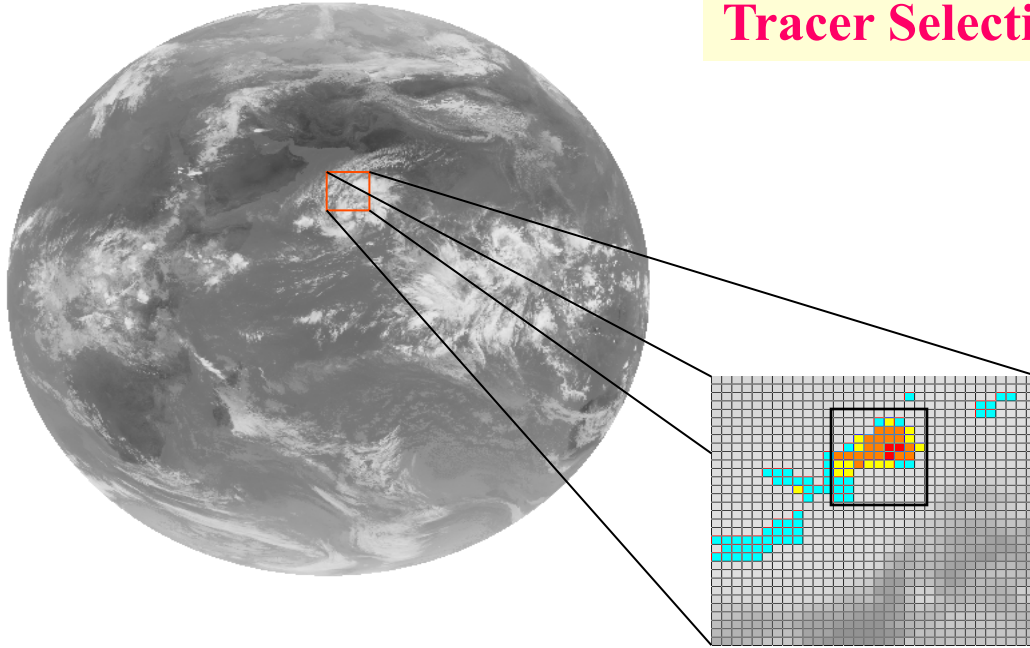
- ❖ Dominant Seasonal Forcing cause the winds to be highly persistent.
- ❖ **Minimum Decorrelation time scale is observed ~ 4 hours.**
- ❖ Each AMV represents 160km x 160km area and loses out on the fine scale features of winds.
- ❖ AMVs themselves have inherent errors ~ 8m/s and 4m/s in mean vector difference for the upper and lower atmospheric levels (Borde et al , 2010) which may affect the accuracy of the decorrelation time scales.
- ❖ This information can be potentially useful in the optimization issues eg. in assimilation , validation and retrieval.
- ❖ Can be used for the Short term empirical prediction.

Recent modification in retrieval algorithm

Image-pair



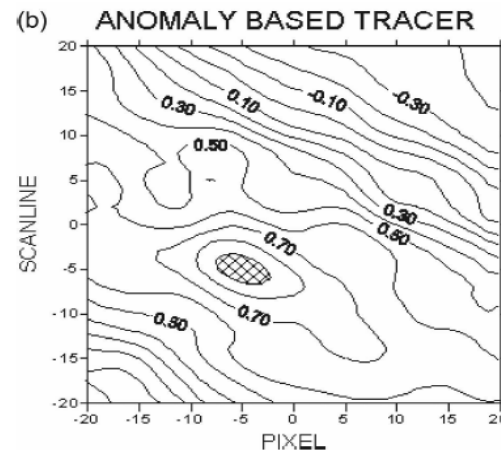
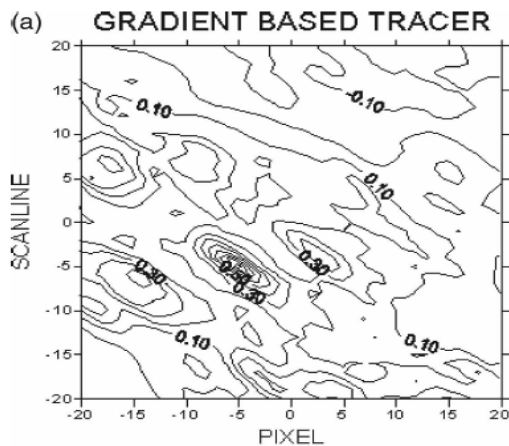
Tracer Selection method



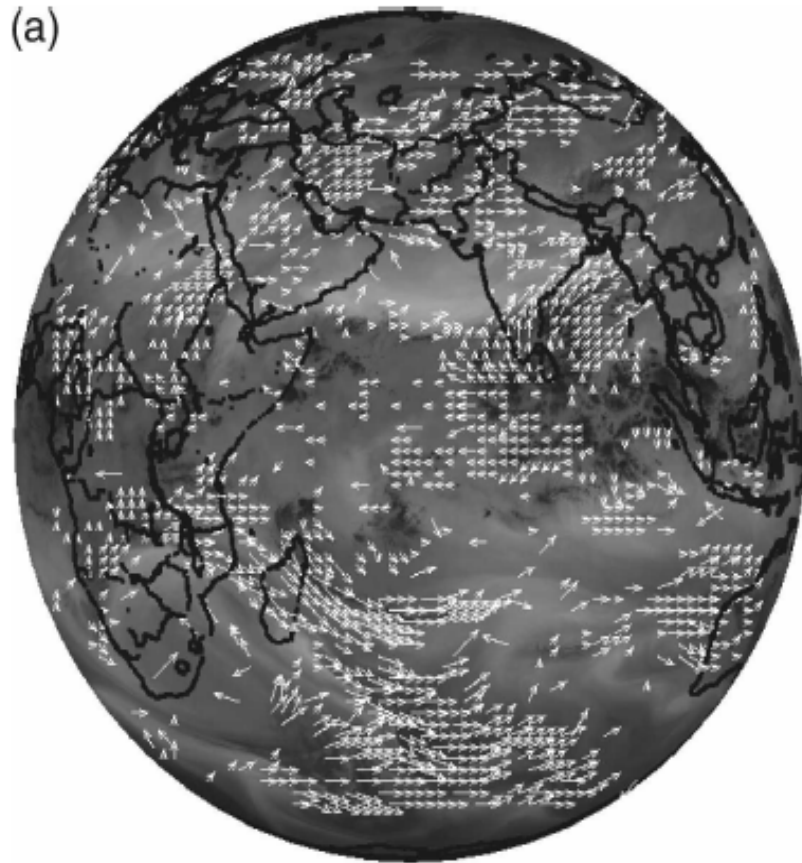
- The tracer selection method based on **local image anomaly**. This method results in **smooth tracer fields** compared to standard “bi-directional gradient” method, particularly in WV images. This can **reduce tracking errors**.

Tracer are selected in 20 x 20 template window

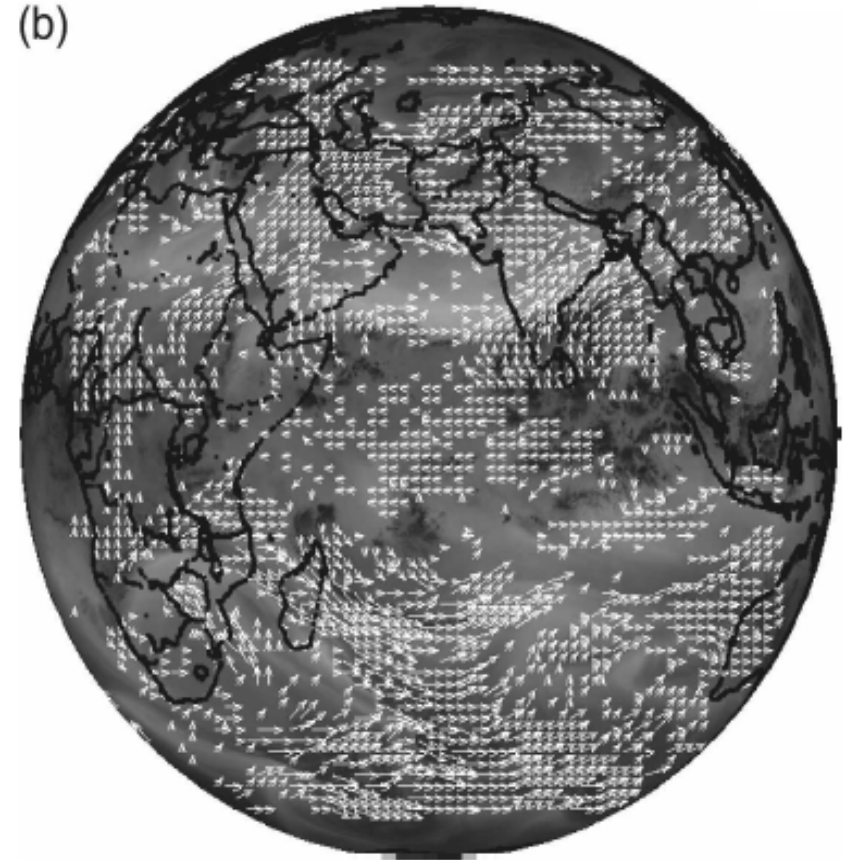
$$\text{Local image anomaly} = \sum_i \sum_j [I(i, j) - \bar{I}],$$



Maximum correlation of WV tracers calculated (from a sample template from 01 Oct 2006 00UTC) a) Bidirectional gradient based and b) anomaly based technique



Gradient based tracer

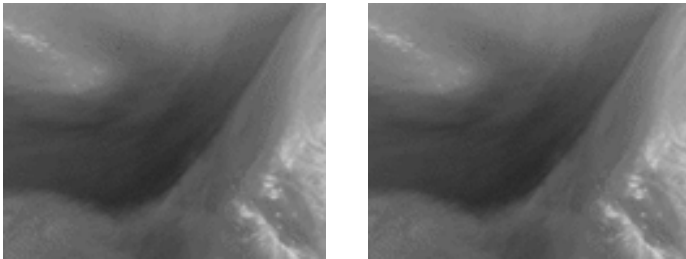


Anomaly based tracer

A sample figure showing the difference of the density of quality-controlled vectors (using identical quality-control criteria) produced by two tracer selection method

Modified Index for Tracking

Maximum match among tracers is determined by using **Nash-Sutcliffe model efficiency** instead of standard Max Cross Correlation (MCC).



Identical Image Pair

One image gradually added with noise

N.S. Fitness index has significantly higher sensitivity to noise compared to MCC thus reducing the chance of picking false target match points.

Nash-Sutcliffe model efficiency

$$E = 1 - \frac{\sum_{i=1}^n (I_t - I_s)^2}{\sum_{i=1}^n (I_t - \bar{I}_t)^2}$$

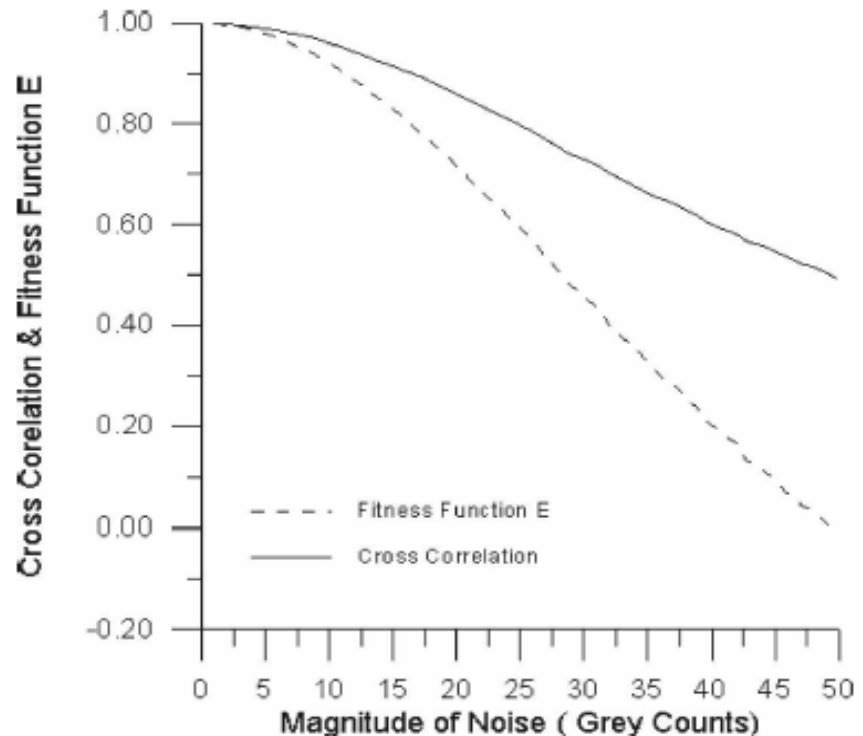


FIG. 3. The values of fitness function E (dashed) and MCC (solid) for varying noise levels.

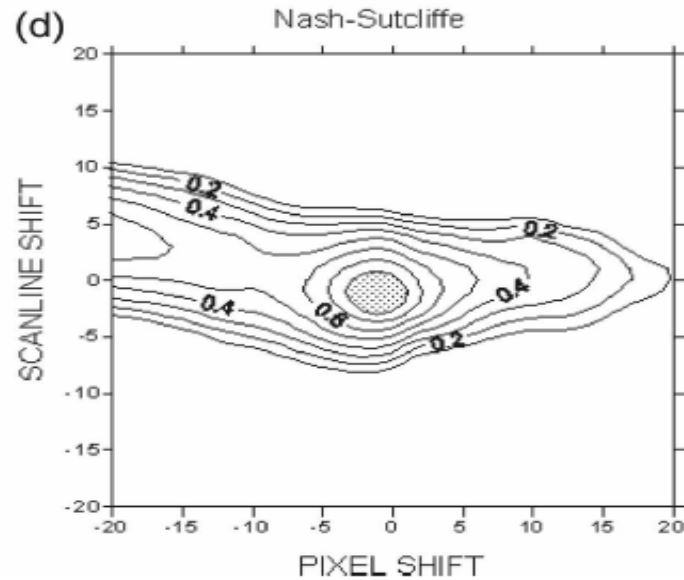
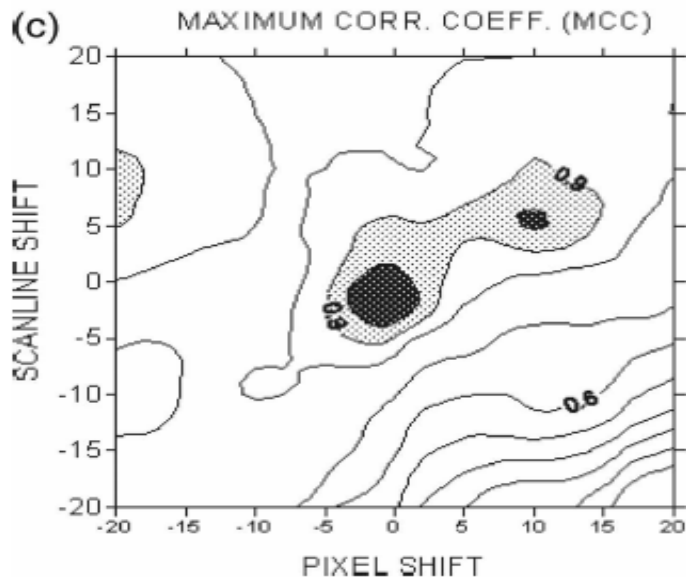
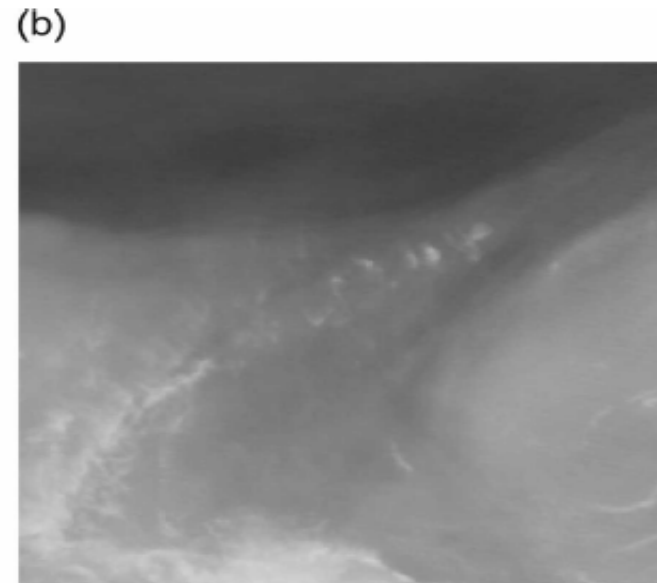
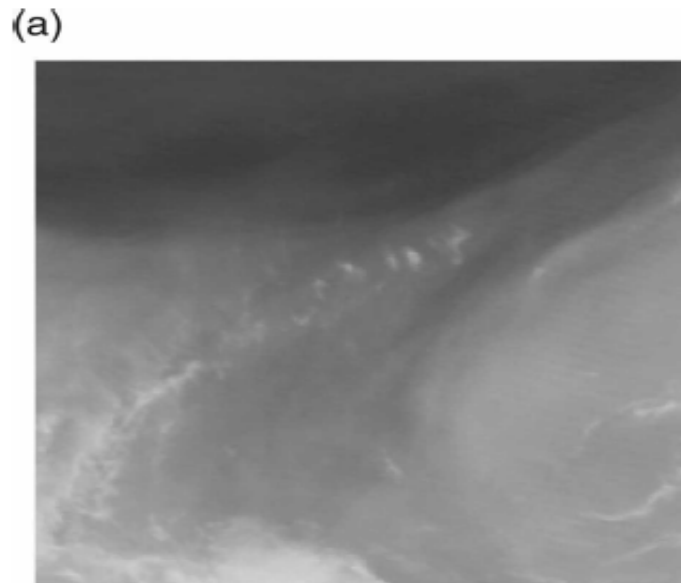
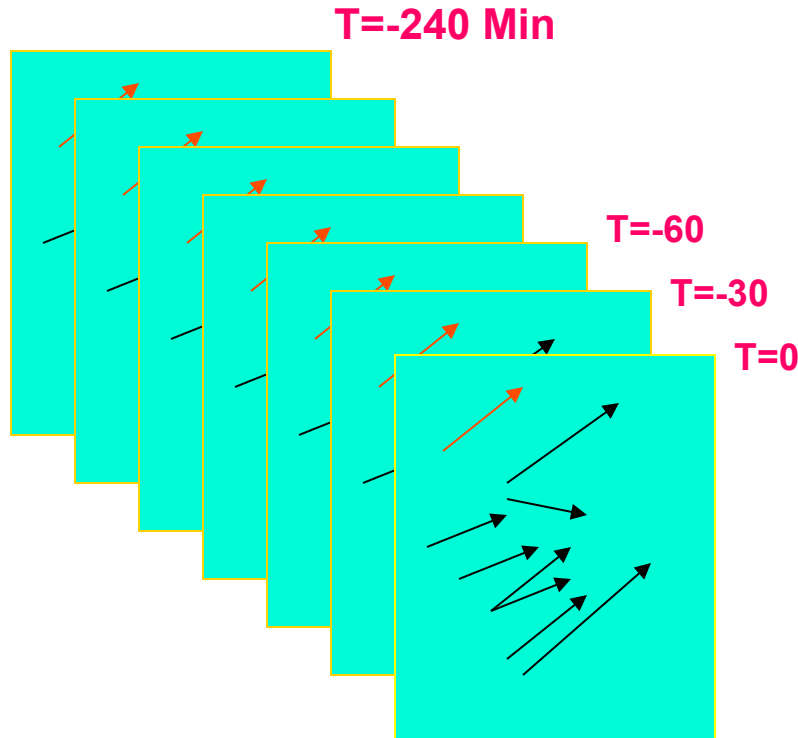


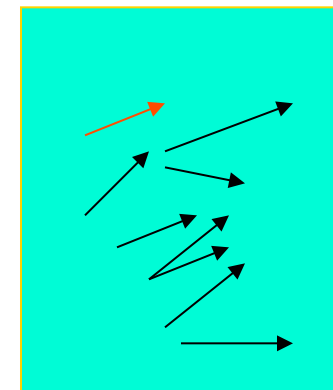
FIG. 2. *Meteosat-5* WV images at (a) 0000 and (b) 0030 UTC. The maximum (c) cross-correlation coefficient and (d) Nash-Sutcliffe efficiency coefficient, calculated during tracking between the image in (a) and the image in (b). Contour levels are -0.2 , 0 , 0.2 , 0.4 , 0.6 , 0.8 , and 0.9 .

Buffer generation and Quality control

- Use of full disc image is replaced with sector generated image with improved registration and fixed lat/lon co-ordinate.
- Take advantage of using multiple 30-min images, rather than traditional 3 images.



Only 1 set retrieved
At a time



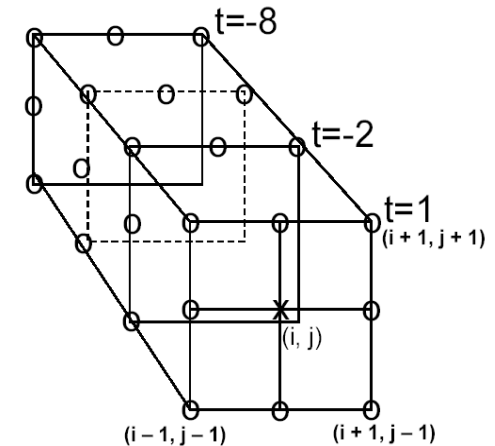
Vector-1

.....All the singular retrievals during past N-hours provide support

?? About the life cycle of cloud.....

Combined Vector Selection and Quality Check

- For every new vector under consideration, (from current image-pair), its vector difference from the buffer is computed as well as in 3×3 neighborhood, provided, the vectors to be compared show similar BT characteristics (e.g. similar levels).



- Each vector difference (magnitude of complex number) is weighted according to distance and time difference from the current vector.
- If the difference of top 30% weighted differences is less than 1.1-pixel, the new vector is accepted, otherwise it is rejected.

Salient Features of this modification

- Utilizes information buffer from past 4-hours for support.
- No thresholds assumed for land/cloud discrimination, making the algorithms more adaptable and dynamic in nature.
- Computationally Faster than previous version.

Height assignment - Collaboration with CIMSS/SSEC



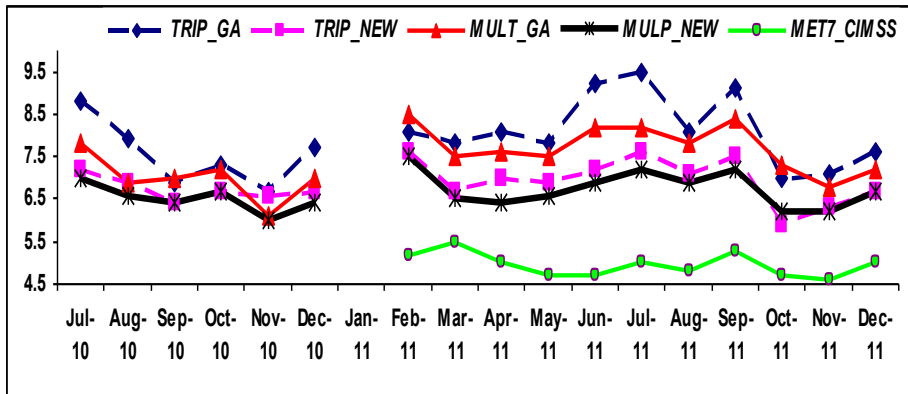
- Infrared window technique (WIN).
- H₂O Intercept Method. (Nieman et. al 1993)
- Cloud Base Method (BASE). (LeMarshall et. al 1993)
- Few gross error checks.

RMSVD

Validation statistics with radiosonde

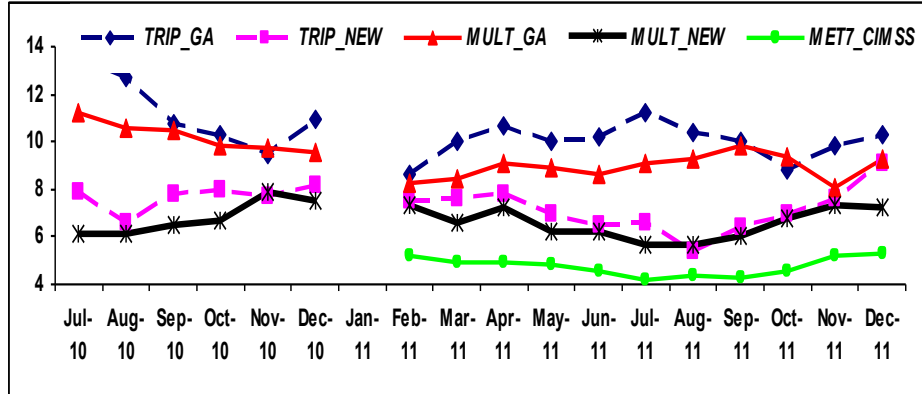
IR Winds

HIGH-Level



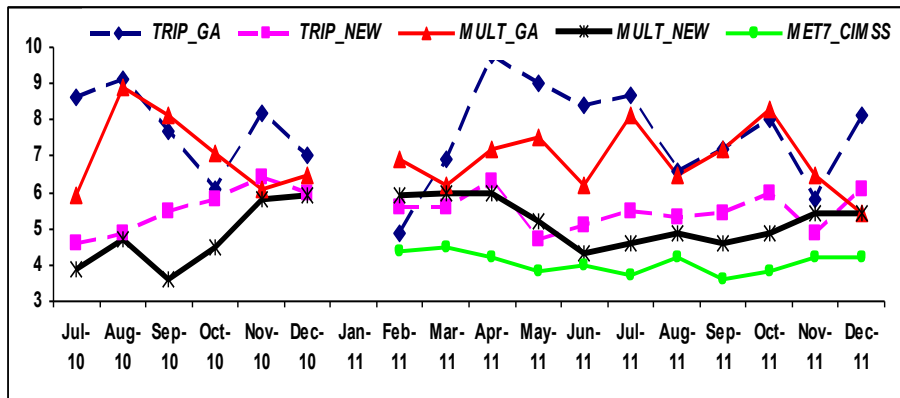
IR Winds

MID-Level



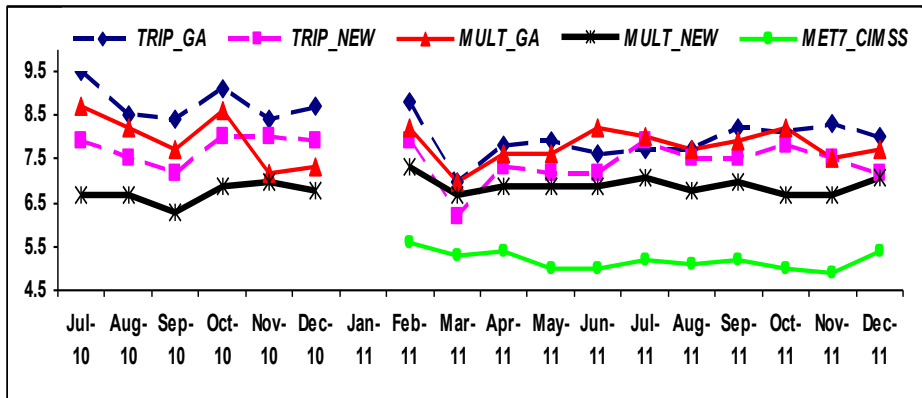
IR Winds

LOW-Level



WV Winds

HIGH-Level



Statistics: Average of all 17 months

| IR-HIGH | | | | |
|---------|---------|----------|----------|-------|
| | TRIP GA | TRIP NEW | MULT NEW | CIMSS |
| MVD | 7.6 | 6.7 | 6.4 | 4.7 |
| RMSVD | 7.9 | 6.8 | 6.6 | 4.9 |
| SD | 1.8 | 1.3 | 1.6 | 1.2 |
| BIAS | -0.8 | -0.7 | -1.4 | -0.9 |
| SP | 15.2 | 16 | 17.7 | 16.3 |
| NC | 4043 | 4671 | 4609 | 4175 |
| IR-MID | | | | |
| MVD | 9.9 | 7 | 6.2 | 4.4 |
| RMSVD | 10.4 | 7.3 | 6.6 | 4.7 |
| SD | 2.9 | 1.6 | 2.1 | 1.2 |
| BIAS | -5.0 | -1.8 | -0.7 | -1.2 |
| SP | 11.3 | 13.2 | 14.6 | 13.7 |
| NC | 883 | 857 | 842 | 848 |
| IR-LOW | | | | |
| MVD | 7.3 | 5.4 | 4.8 | 3.9 |
| RMSVD | 7.6 | 5.5 | 5 | 4.0 |
| SD | 0.9 | 0.4 | 0.9 | 0.4 |
| BIAS | -3.1 | -2.4 | -1.0 | -0.8 |
| SP | 7.8 | 7.8 | 8.5 | 8.8 |
| NC | 93 | 86 | 175 | 160 |
| WV-HIGH | | | | |
| MVD | 7.6 | 7.1 | 6.4 | 4.9 |
| RMSVD | 8.2 | 7.5 | 6.7 | 5.1 |
| SD | 2.8 | 2.15 | 1.8 | 1.6 |
| BIAS | -1.2 | -0.4 | -0.4 | -0.3 |
| SP | 18.3 | 18.4 | 17.9 | 17.2 |
| NC | 2873 | 2985 | 4766 | 3706 |

Filters:

Hor. Dist. < 110 Km
 Vert. Dist. < 25 hPa
 Speed Diff. < 30 m/s
 Dir Diff. < 60 deg
 AMV Speed > 2.5 m/s

Approx. 7-9% of total collocations are affected due to these filters.

Spatial error analysis and density plot

IR winds HIGH – Level Period: DEC2011

Bias

TRIP_GA

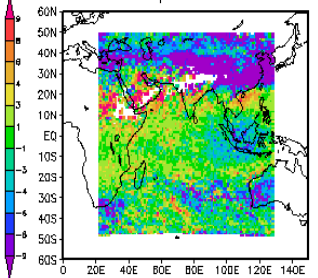
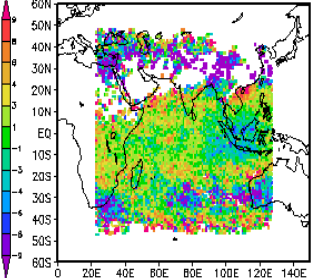
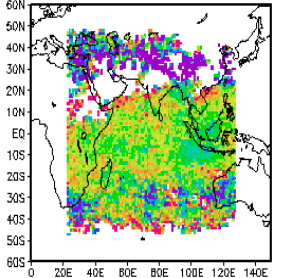
TRIP_NEW

MULT_NEW

O-B Speed Bias

O-B Speed Bias

O-B Speed Bias



MVD

TRIP_GA

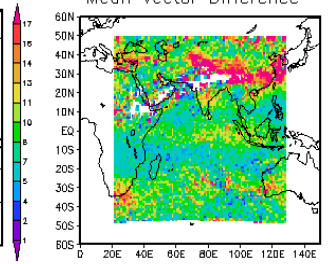
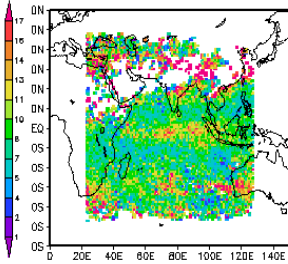
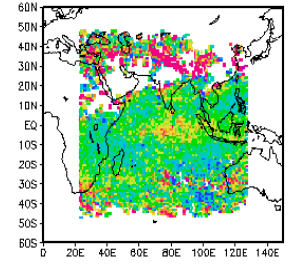
TRIP_NEW

MULT_NEW

Mean Vector Difference

Mean Vector Difference

Mean Vector Difference

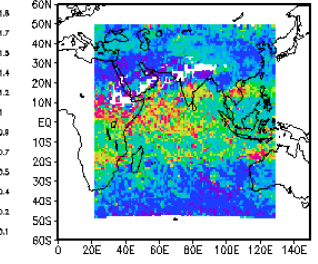
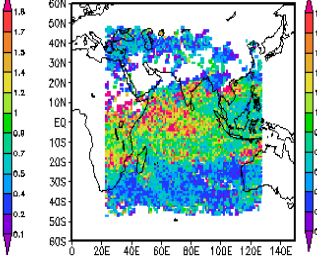
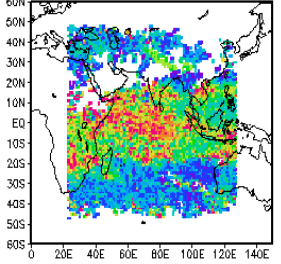


RMSVD (Normalized)

Normalised RMS Vector Difference

Normalised RMS Vector Difference

Normalised RMS Vector Difference

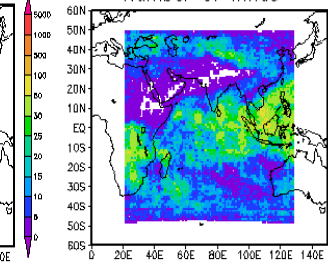
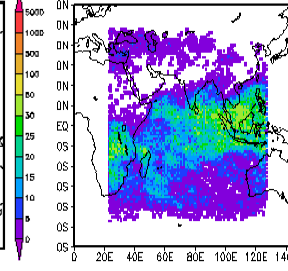
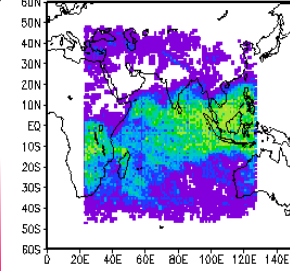


Collocations

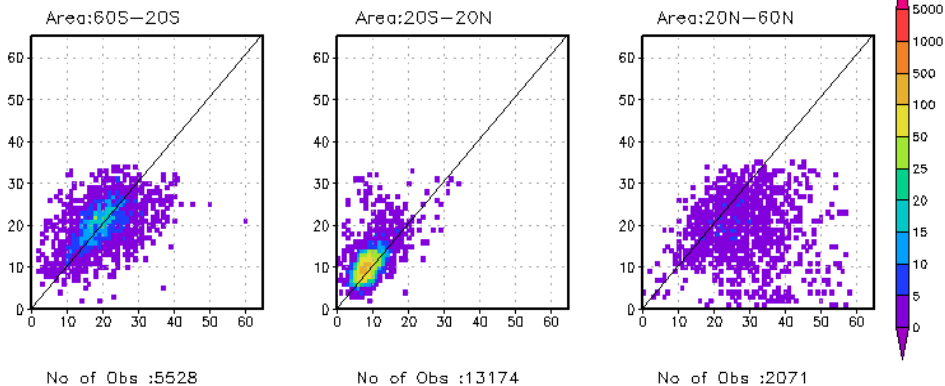
Number of Winds

Number of Winds

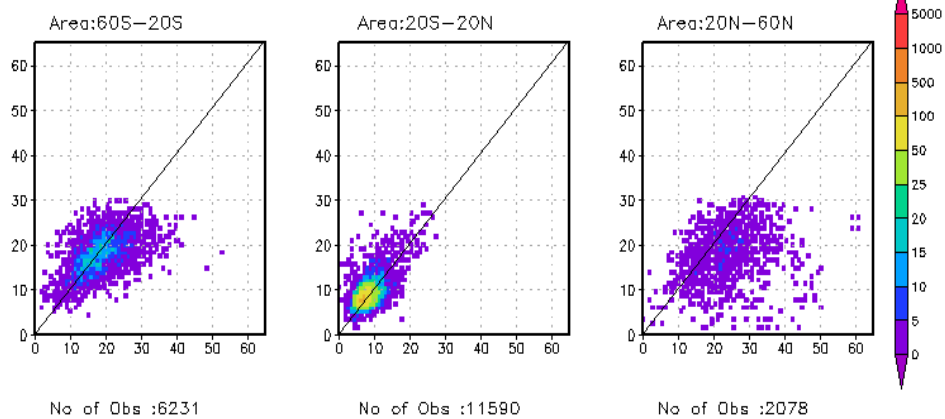
Number of Winds



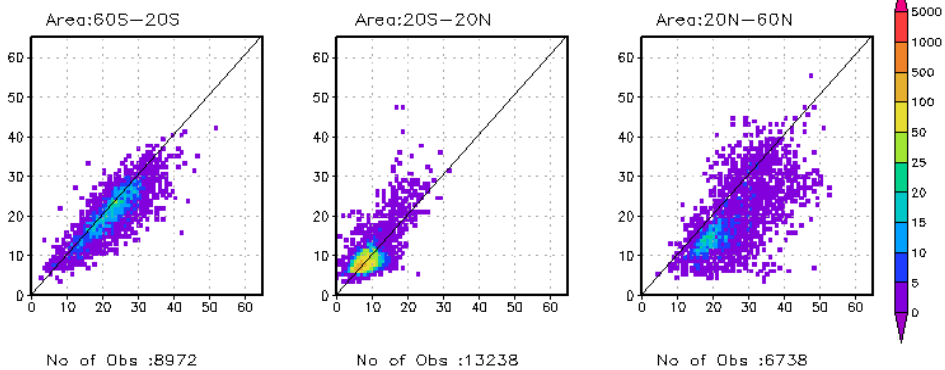
Density plot: IR winds HIGH – Level Period: DEC2011



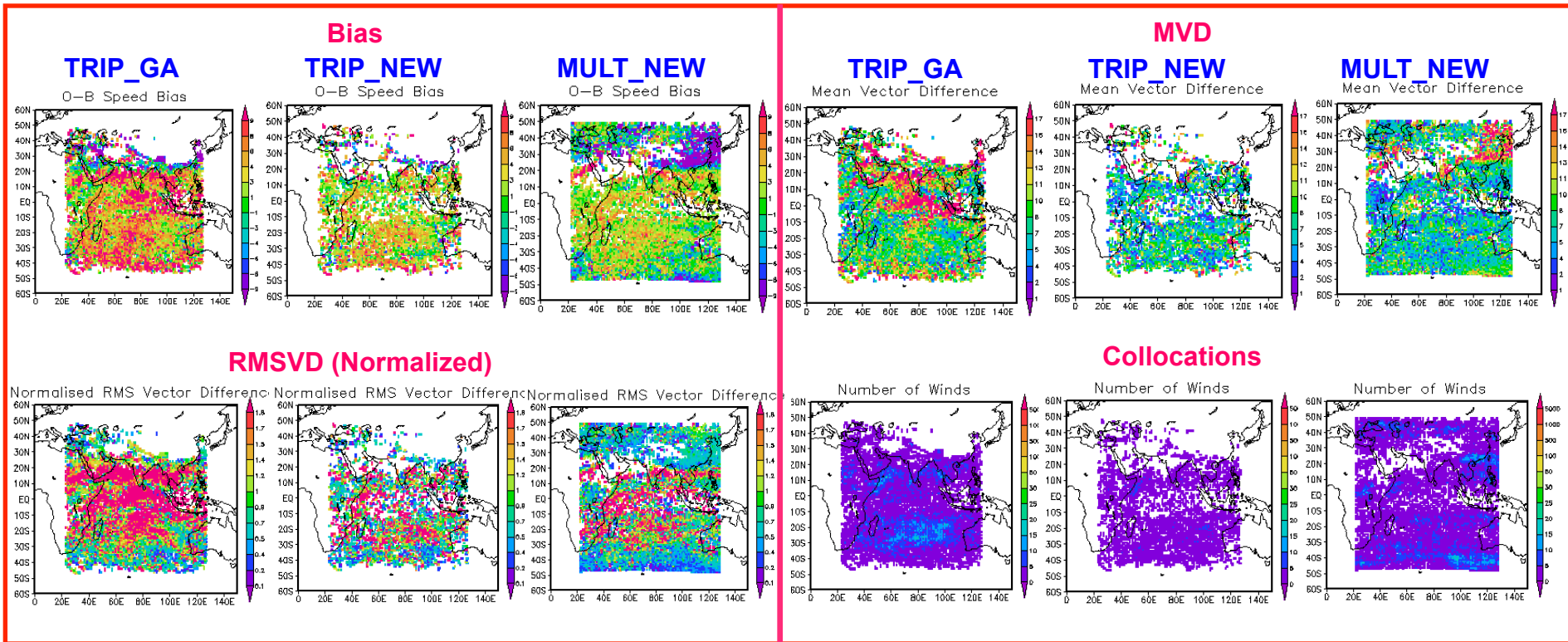
TRIP_GA

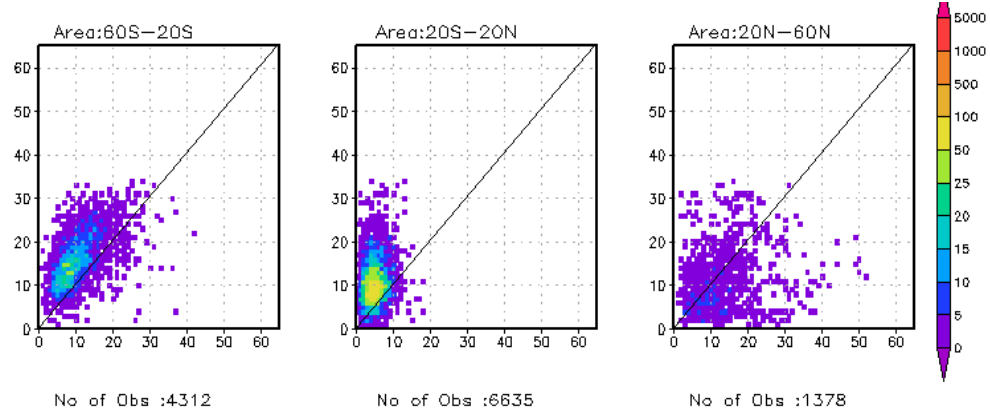


TRIP_NEW

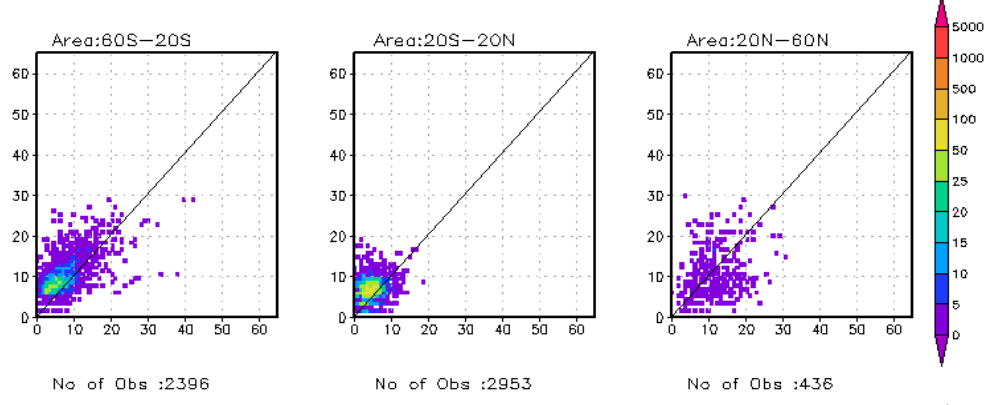


MULT_NEW

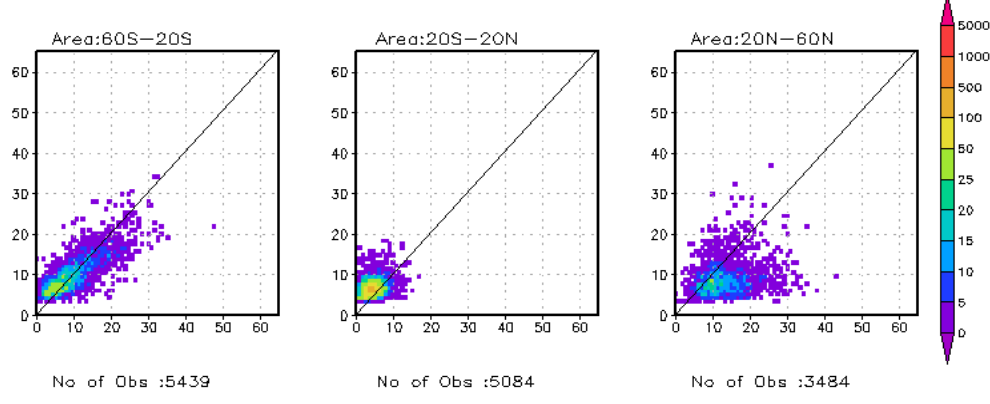




TRIP_GA



TRIP_NEW



MULT_NEW

IR winds LOW – Level

Period: DEC2011

Bias

TRIP_GA

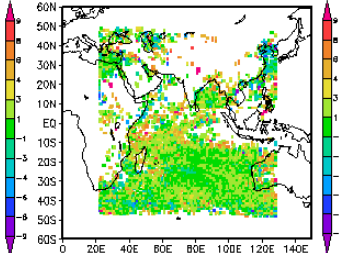
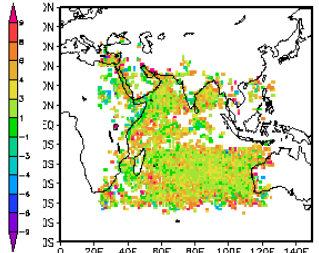
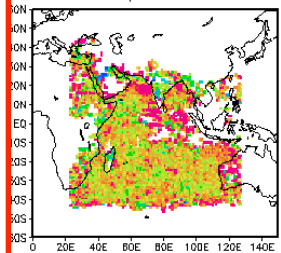
TRIP_NEW

MULT_NEW

O-B Speed Bias

O-B Speed Bias

O-B Speed Bias



MVD

TRIP_GA

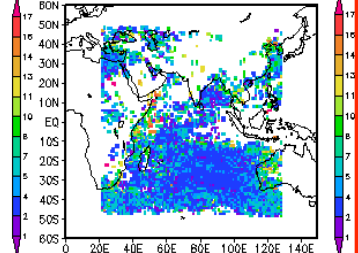
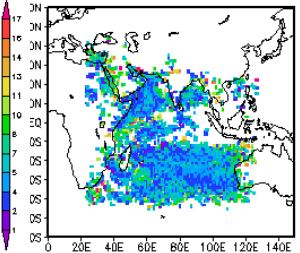
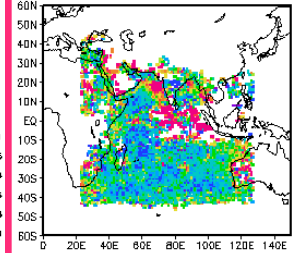
TRIP_NEW

MULT_NEW

Mean Vector Difference

Mean Vector Difference

Mean Vector Difference

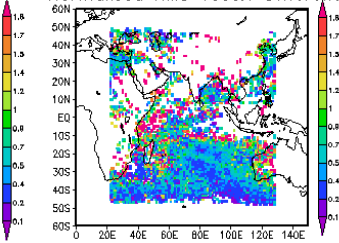
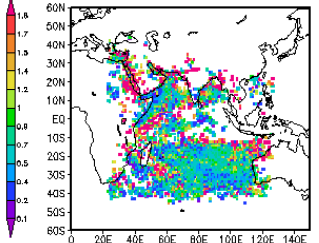
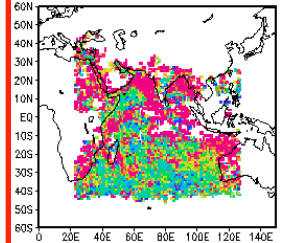


RMSVD (Normalized)

Normalised RMS Vector Difference

Normalised RMS Vector Difference

Normalised RMS Vector Difference

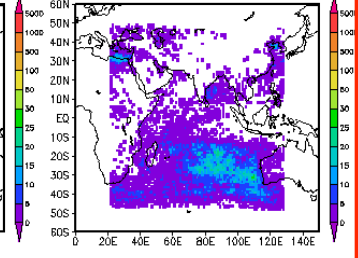
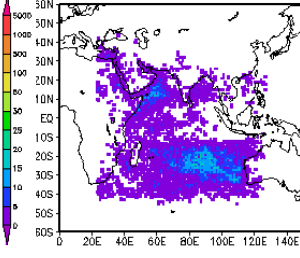
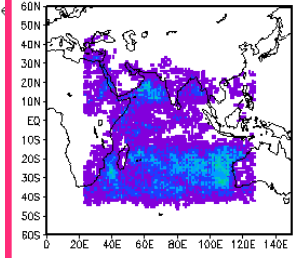


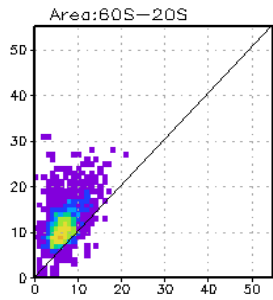
Collocations

Number of Winds

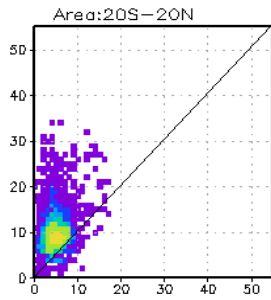
Number of Winds

Number of Winds

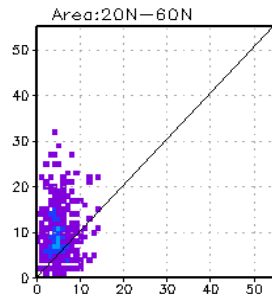




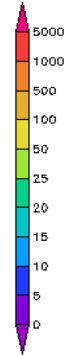
No of Obs :3940



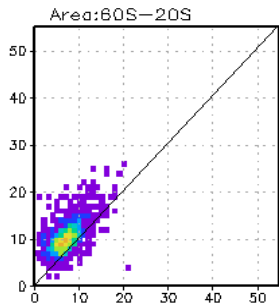
No of Obs :3749



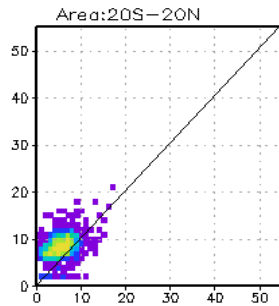
No of Obs :745



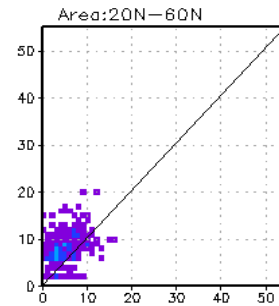
TRIP_GA



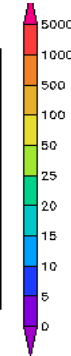
No of Obs :2758



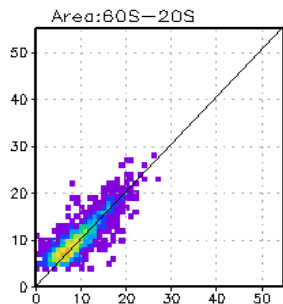
No of Obs :2637



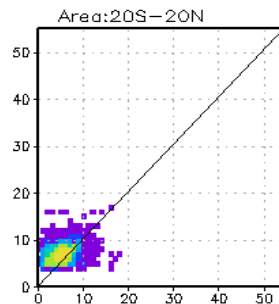
No of Obs :433



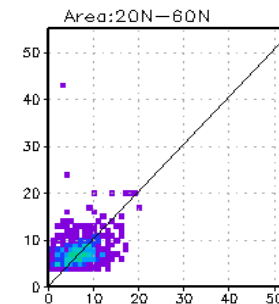
TRIP_NEW



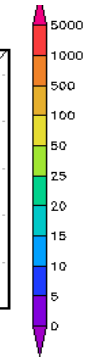
No of Obs :4321



No of Obs :2473



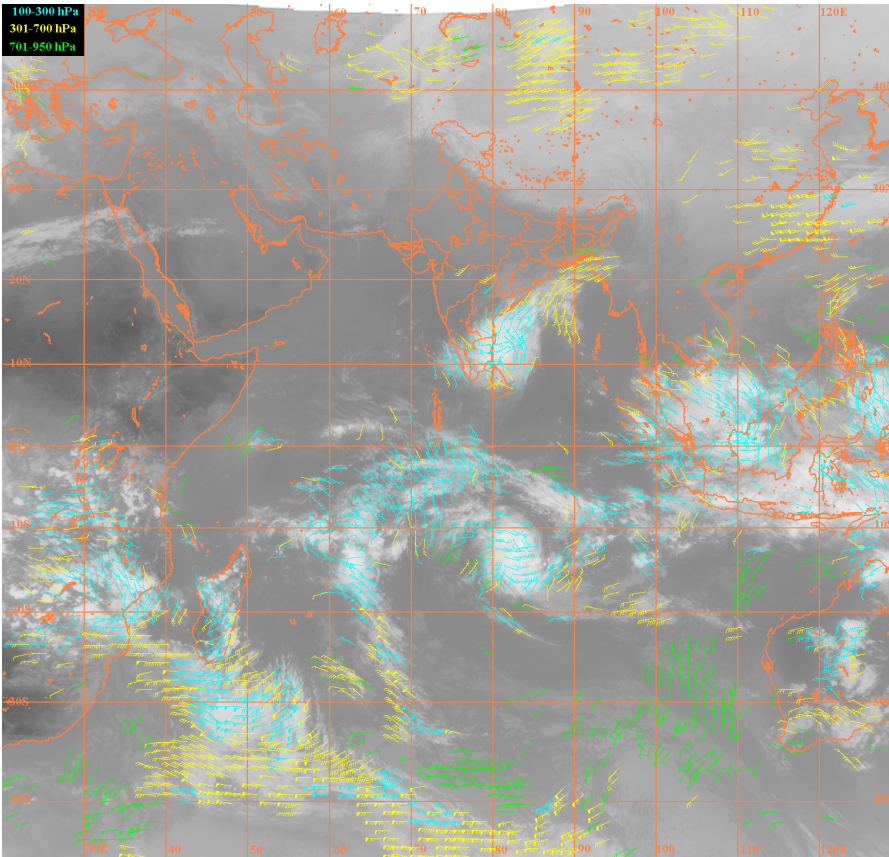
No of Obs :1089



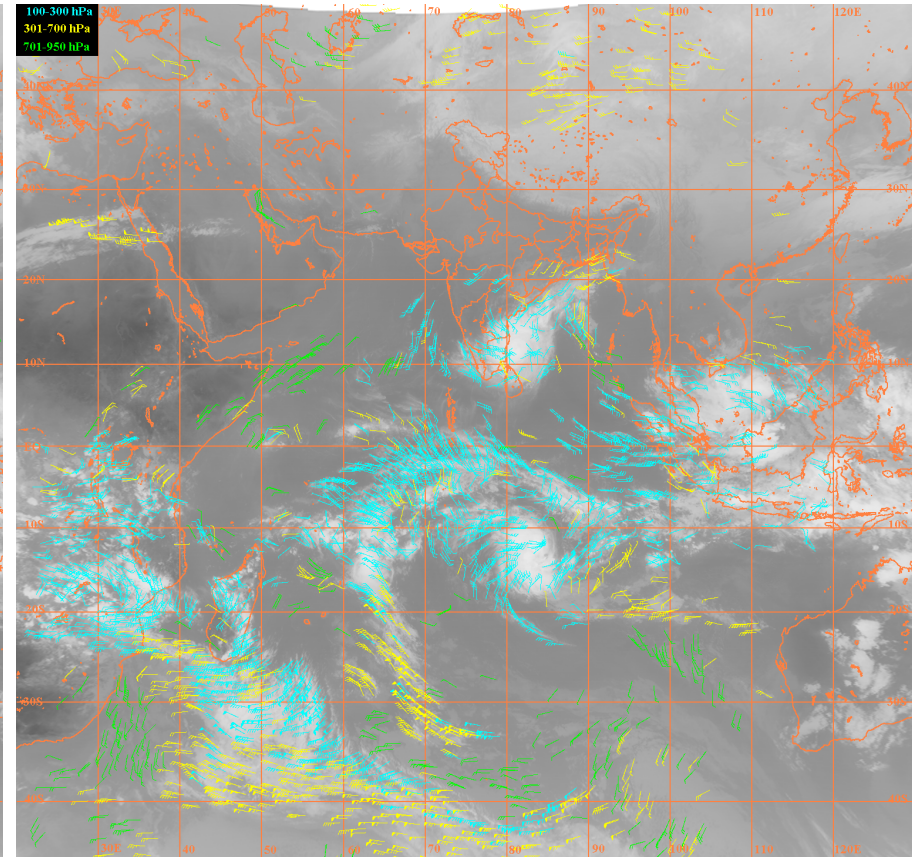
MULT_NEW

29 December 2011 12 UTC: Sample picture cyclone Thane

Kalpana-1-ISRO



Infrared winds



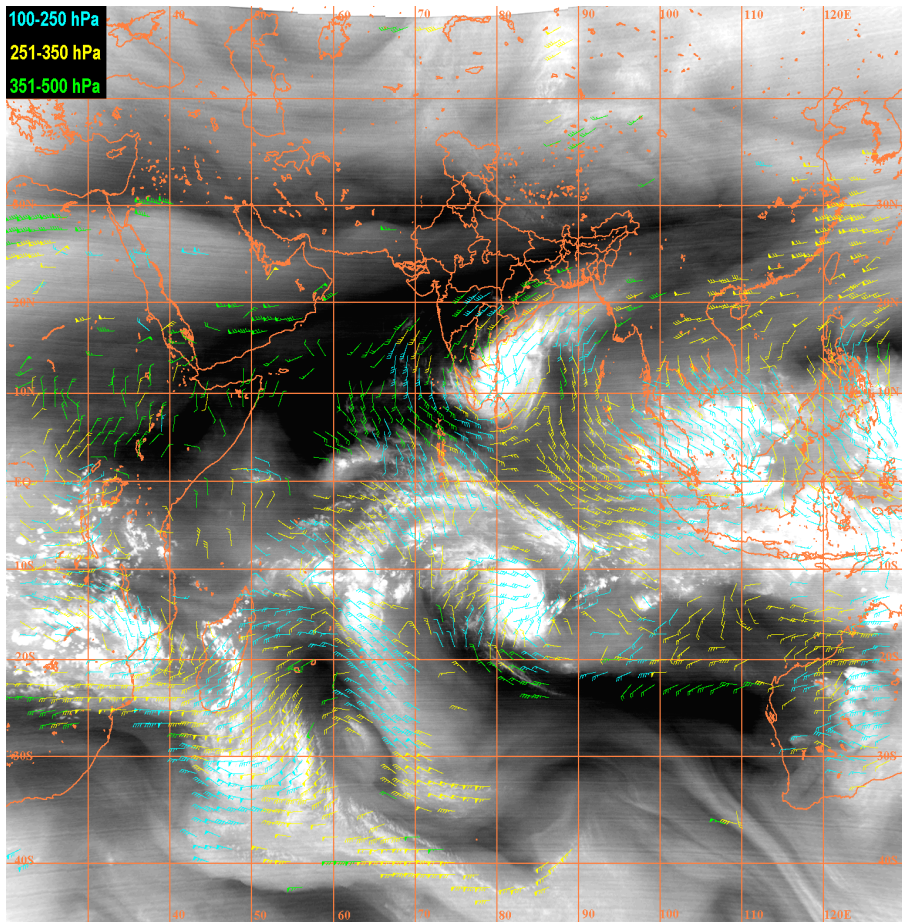
Meteosat7- CIMSS



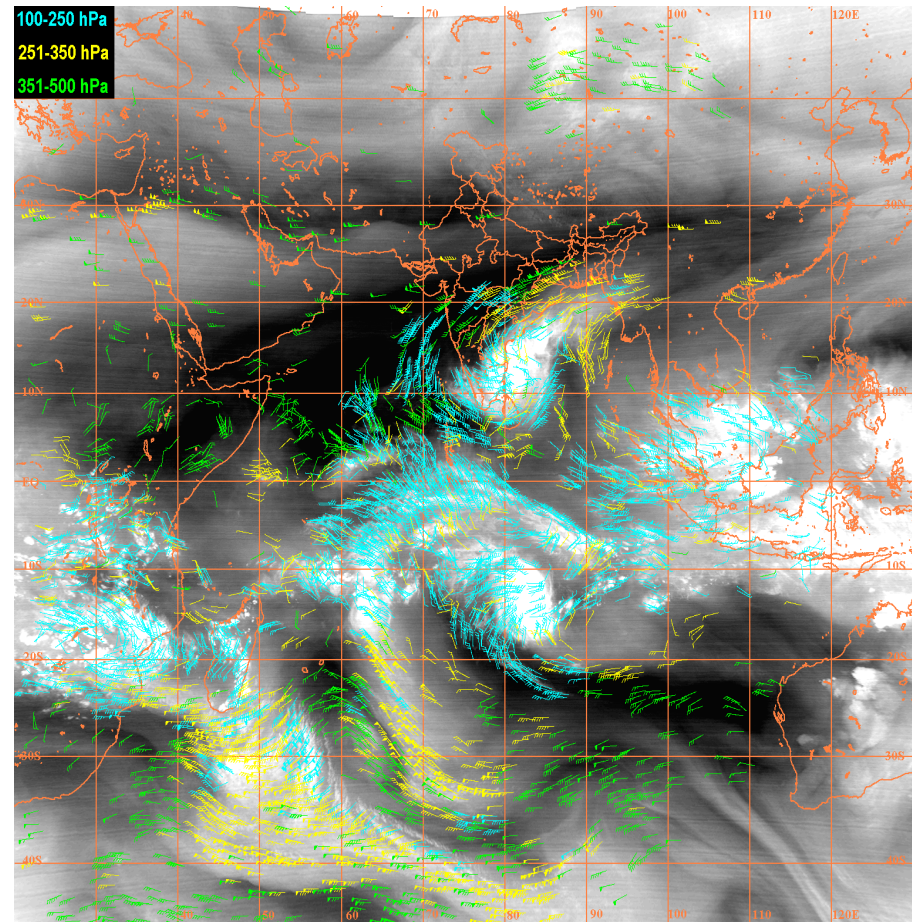
- Thanks to CIMSS for Met7 winds

29 December 2011 12 UTC: Sample picture cyclone Thane

Kalpana-1-ISRO



Water vapor winds



Meteosat7- CIMSS



- Thanks to CIMSS for Met7 winds

Findings from this study

- New algorithm results in **25% more accurate vectors**.
- Higher number of accurate retrievals in low level.
- Captures “meridional flow” at upper level better than triplet algorithm.
- Significant improvement due to new height assignment algorithm.
(RMSVD: high: 7.9 -> 6.8, mid: 10.4 -> 7.3, Low: 7.6 -> 5.5 and WV high: 8.2 -> 7.5)
- However, some positive impact is noticed due to new MULTIPLET algorithm.
(RMSVD: high: 6.8 -> 6.6, mid: 7.3 -> 6.6, Low: 5.5 -> 5.0 and WV high: 7.5 -> 6.7)

Future directions:

- Expect some more improvement with incorporation of auto editor.
- Optimization in tracer selection.
- To retrieve winds from visible and 3.9 μm channel.
- Wish to collaborate with other operational agency for further improvement.

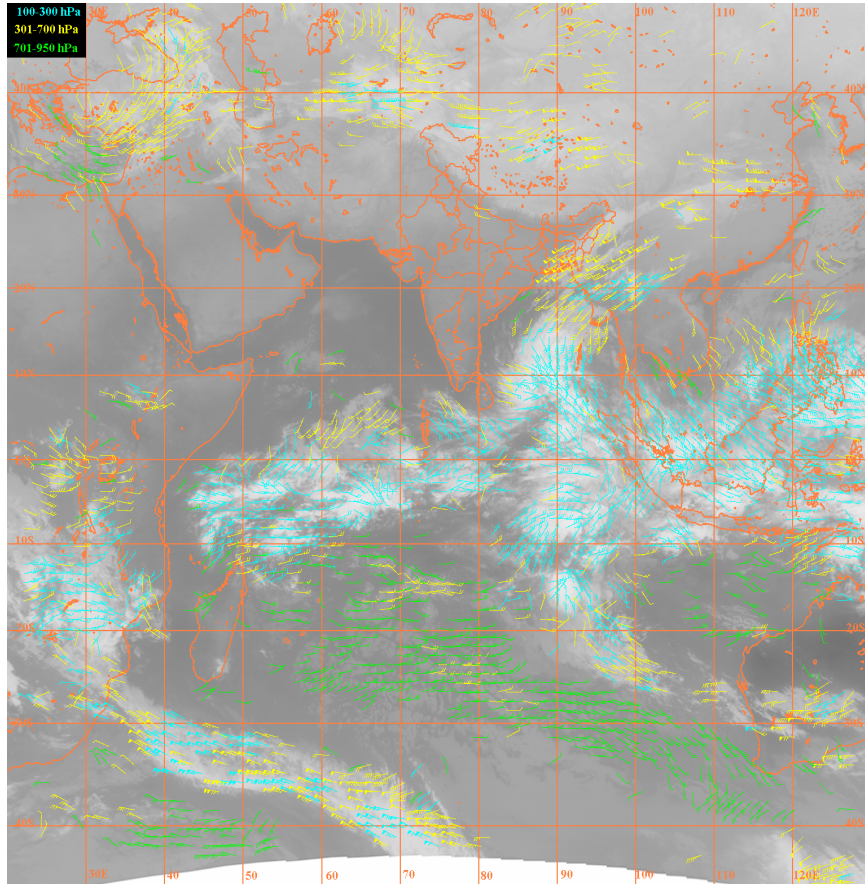
Acknowledgement

- EUMETSAT for financial support for attending 11IWW
- SSEC/CIMSS for scientific collaboration.
- 11IWW organizing committee for all support.
- Director and Associate Director SAC/ISRO for all support.
- EUMETSAT/CIMSS/UKMO for using their AMV products and analysis report.

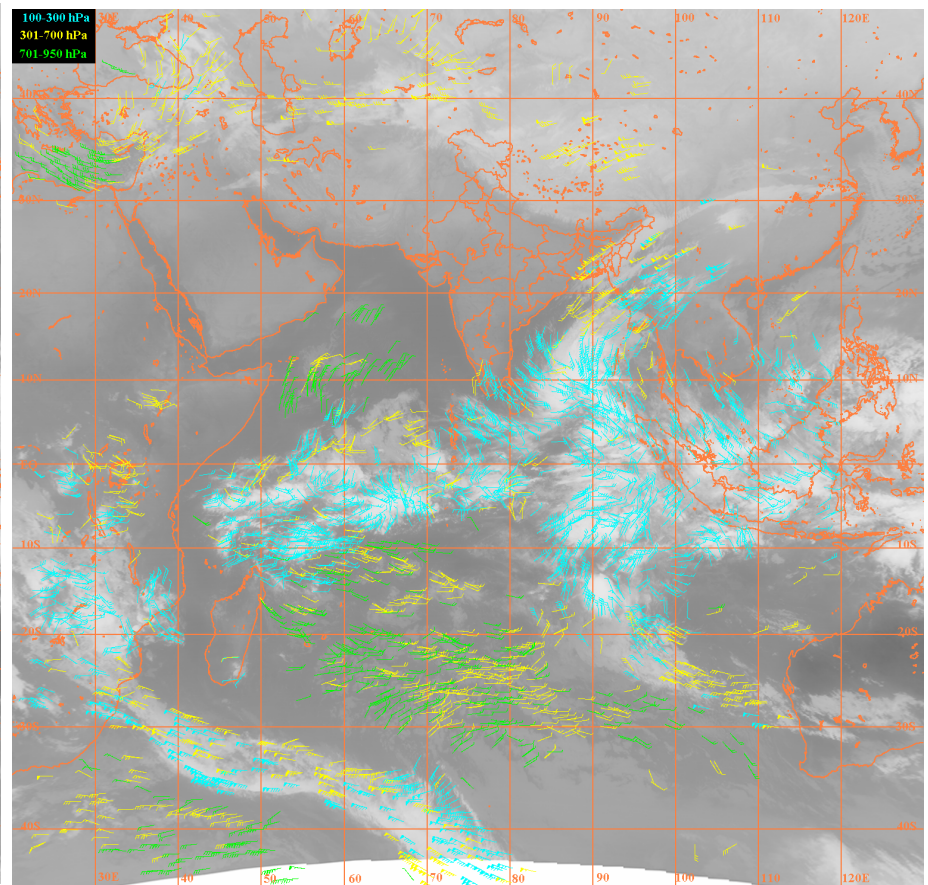
THANK YOU

Cyclone Thane: IR winds - 25th December 2011: 00 UTC

Kalpana-1/ISRO

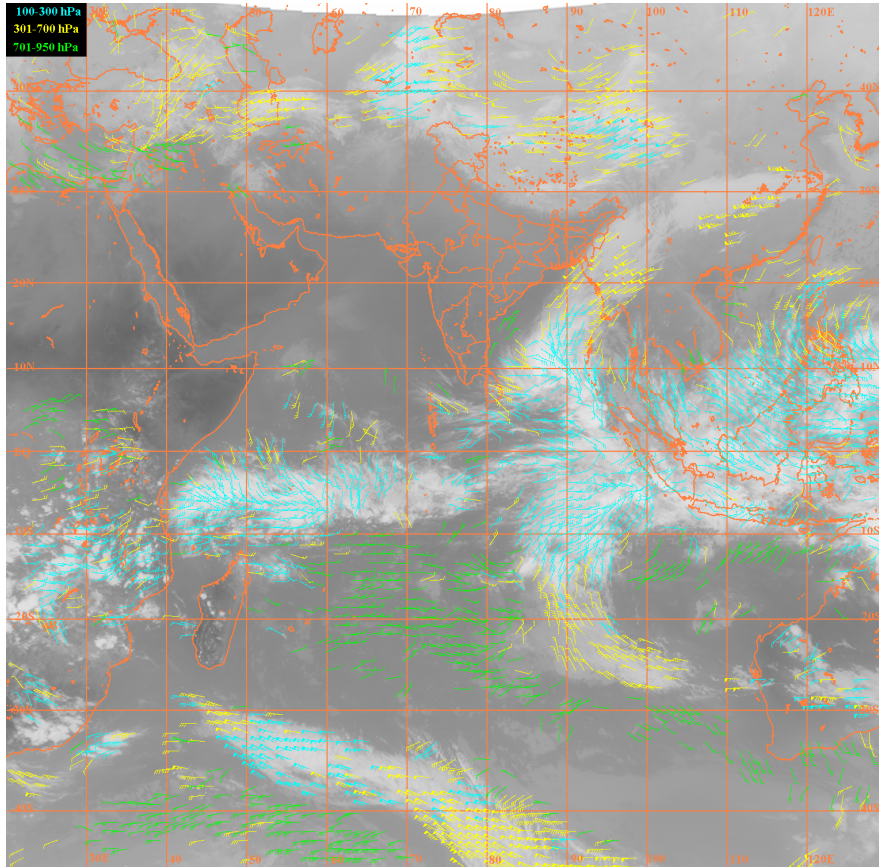


Meteosat7/CIMSS

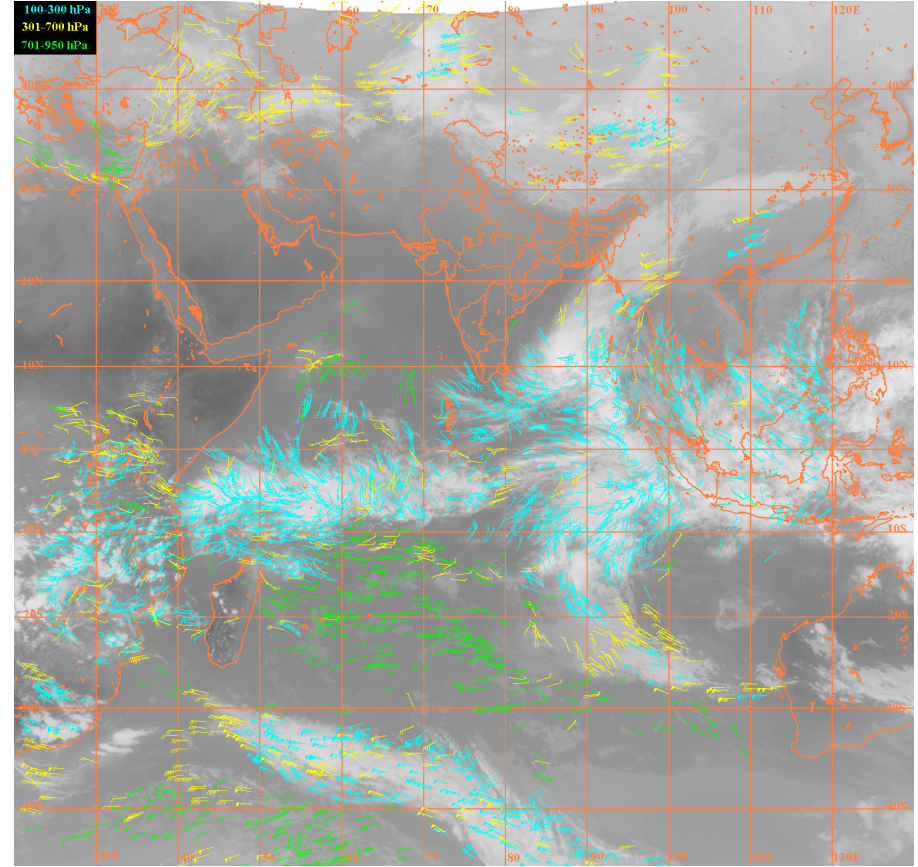


Cyclone Thane: IR winds - 25th December 2011: 12 UTC

Kalpana-1/ISRO

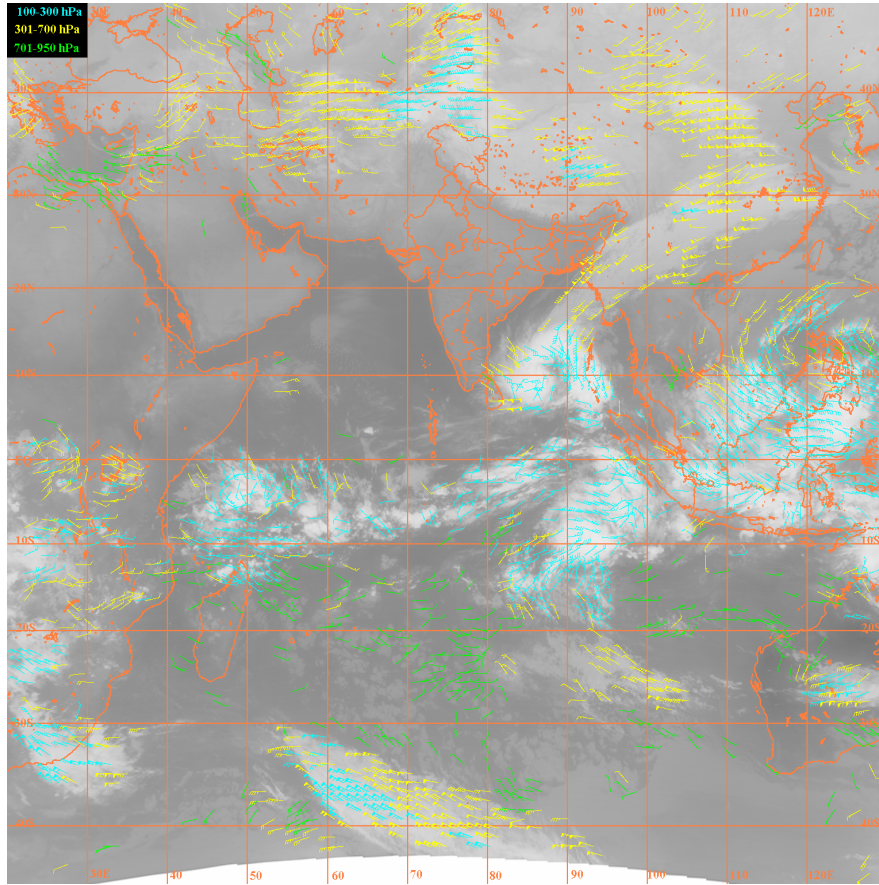


Meteosat7/CIMSS

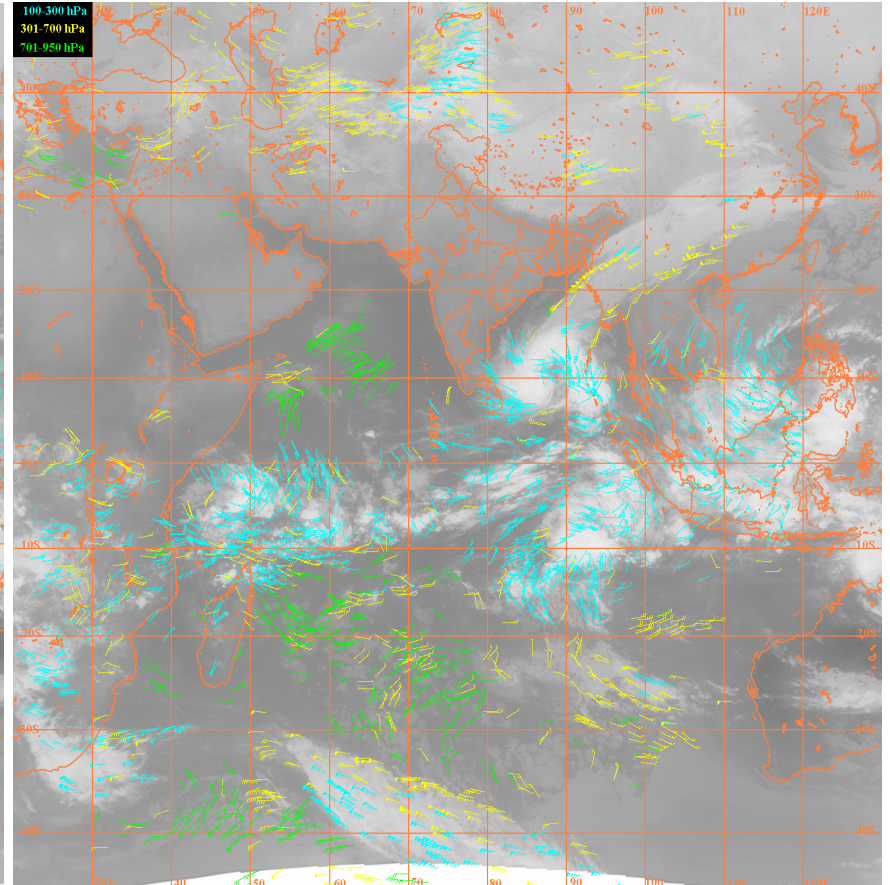


Cyclone Thane: IR winds - 26th December 2011: 00 UTC

Kalpana-1/ISRO

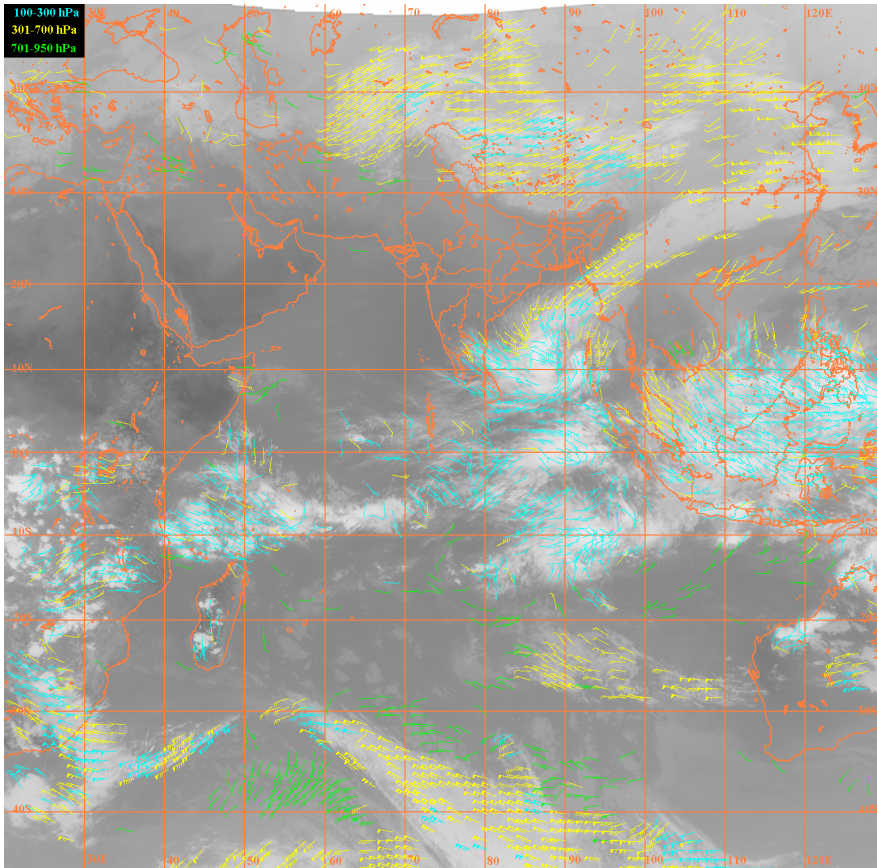


Meteosat7/CIMSS

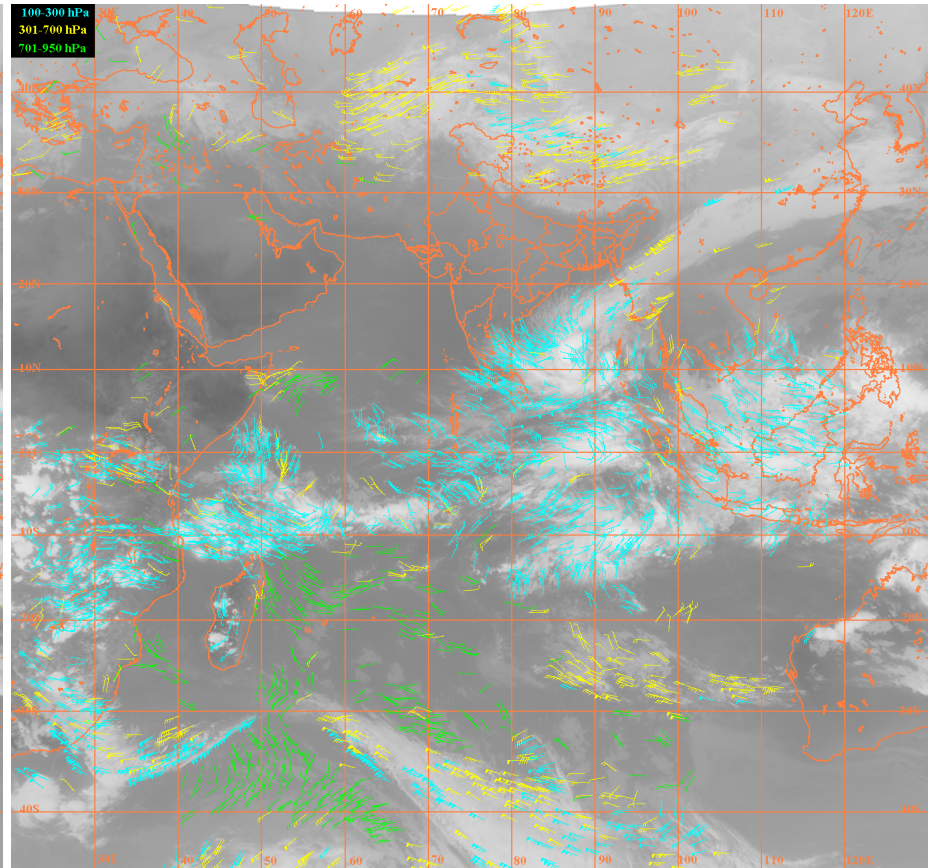


Cyclone Thane: IR winds - 26th December 2011: 12 UTC

Kalpana-1/ISRO

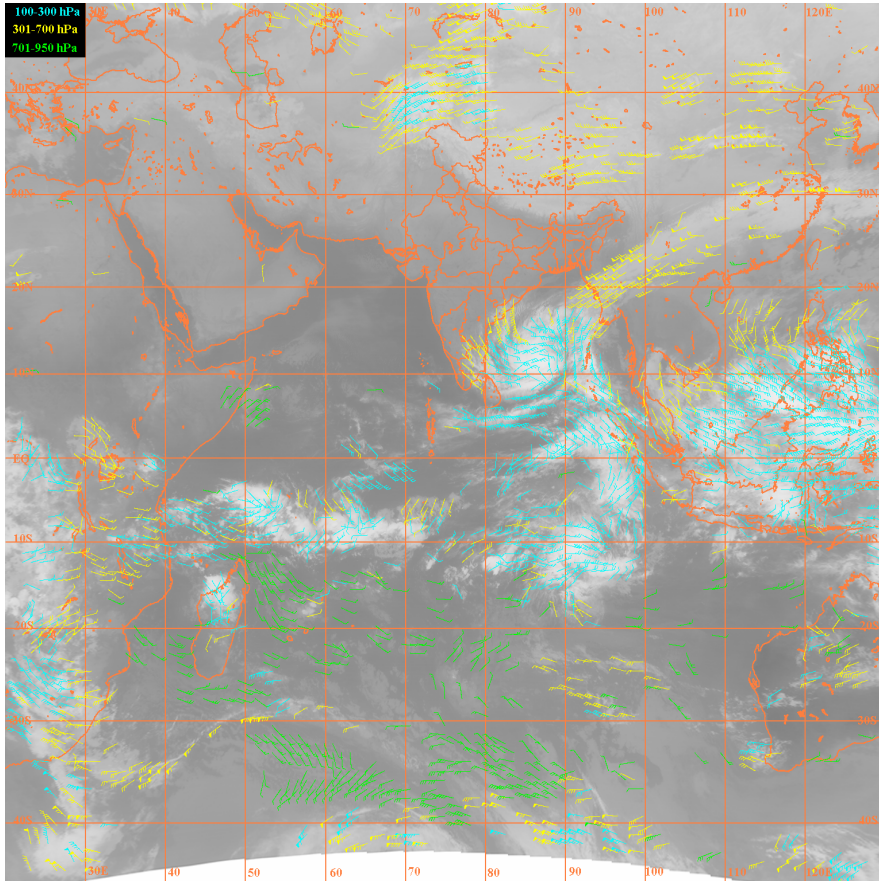


Meteosat7/CIMSS

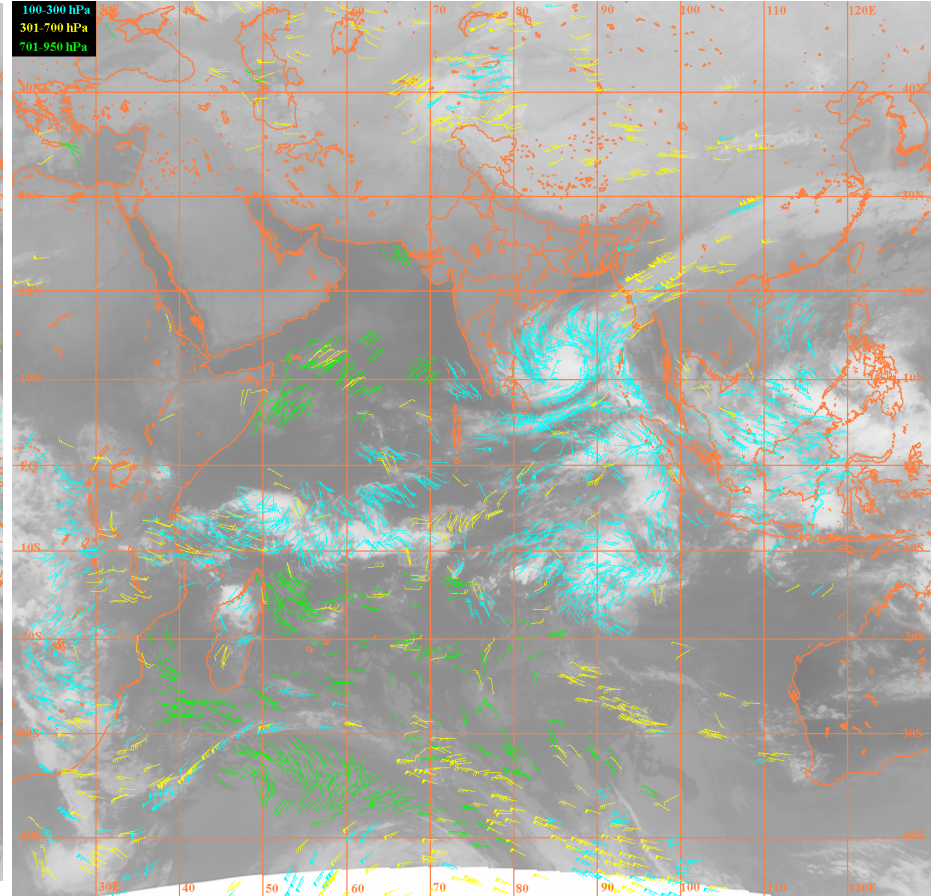


Cyclone Thane: IR winds - 27th December 2011: 00 UTC

Kalpana-1/ISRO

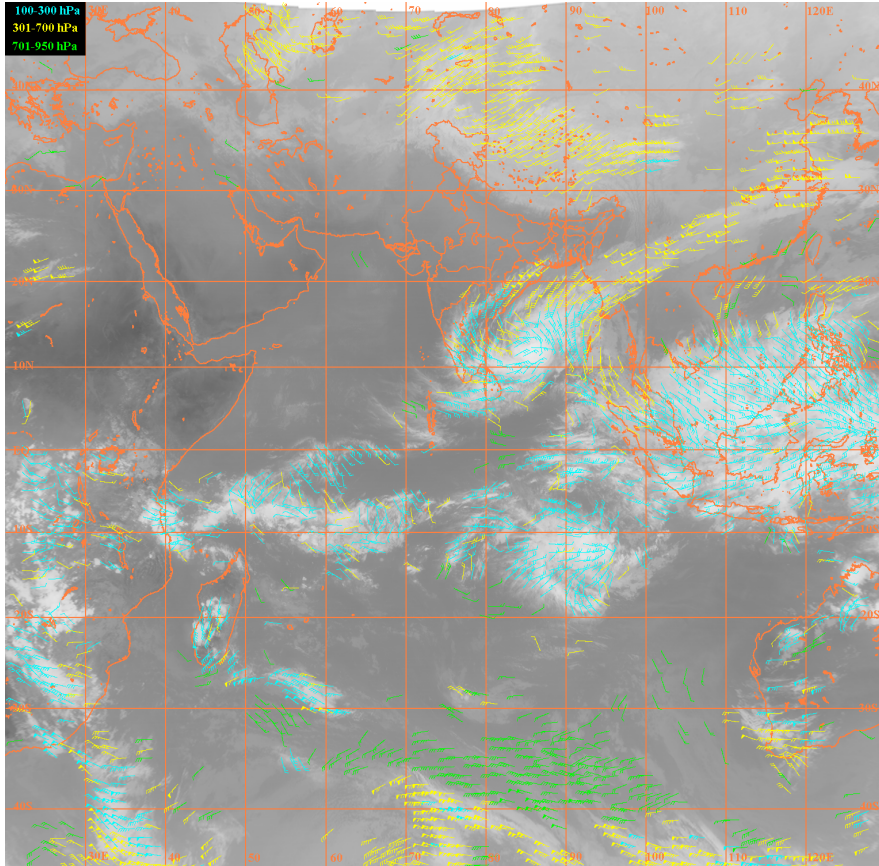


Meteosat7/CIMSS

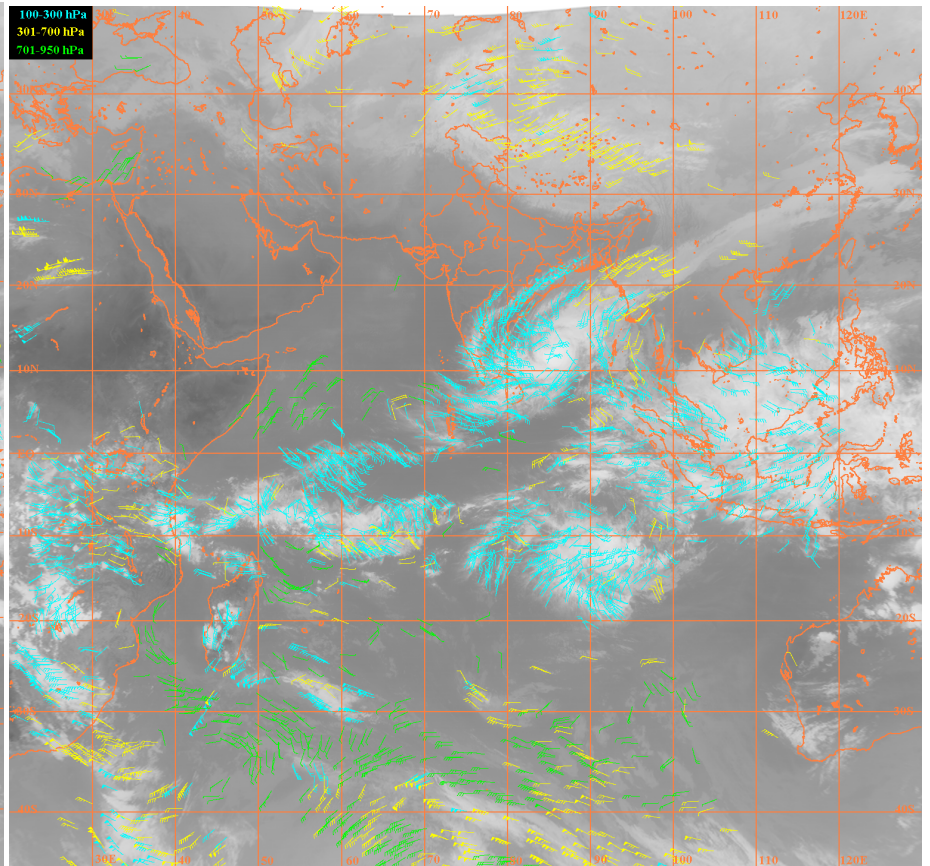


Cyclone Thane: IR winds - 27th December 2011: 12 UTC

Kalpana-1/ISRO

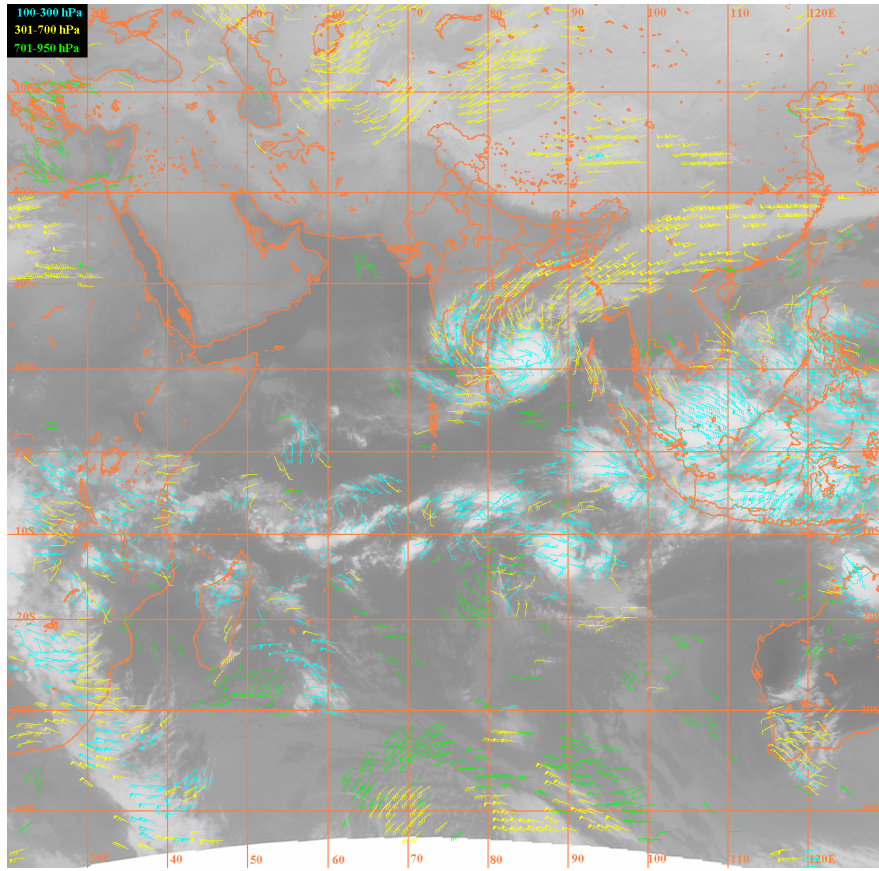


Meteosat7/CIMSS

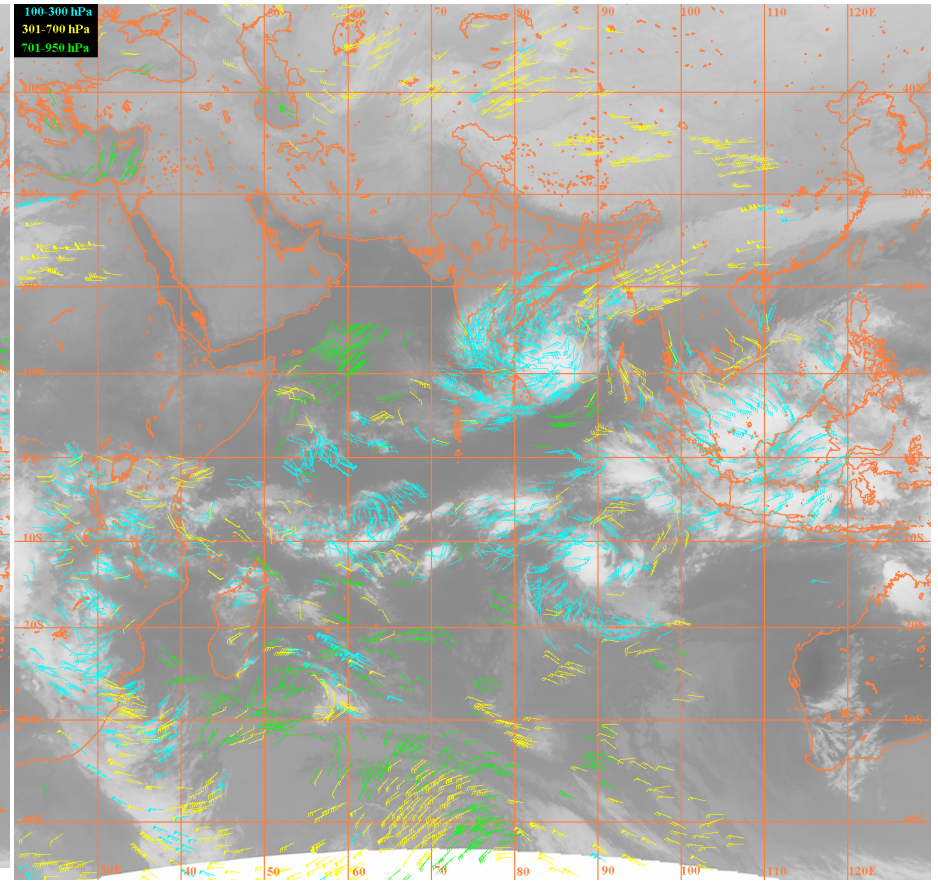


Cyclone Thane: IR winds - 28th December 2011: 00 UTC

Kalpana-1/ISRO

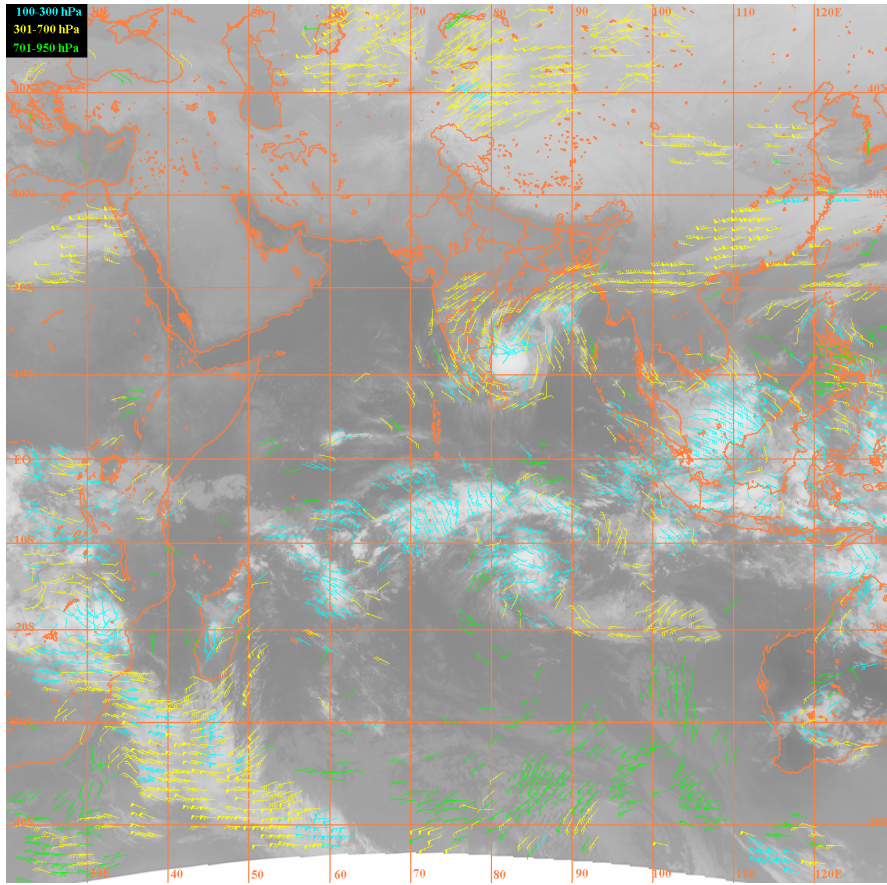


Meteosat7/CIMSS

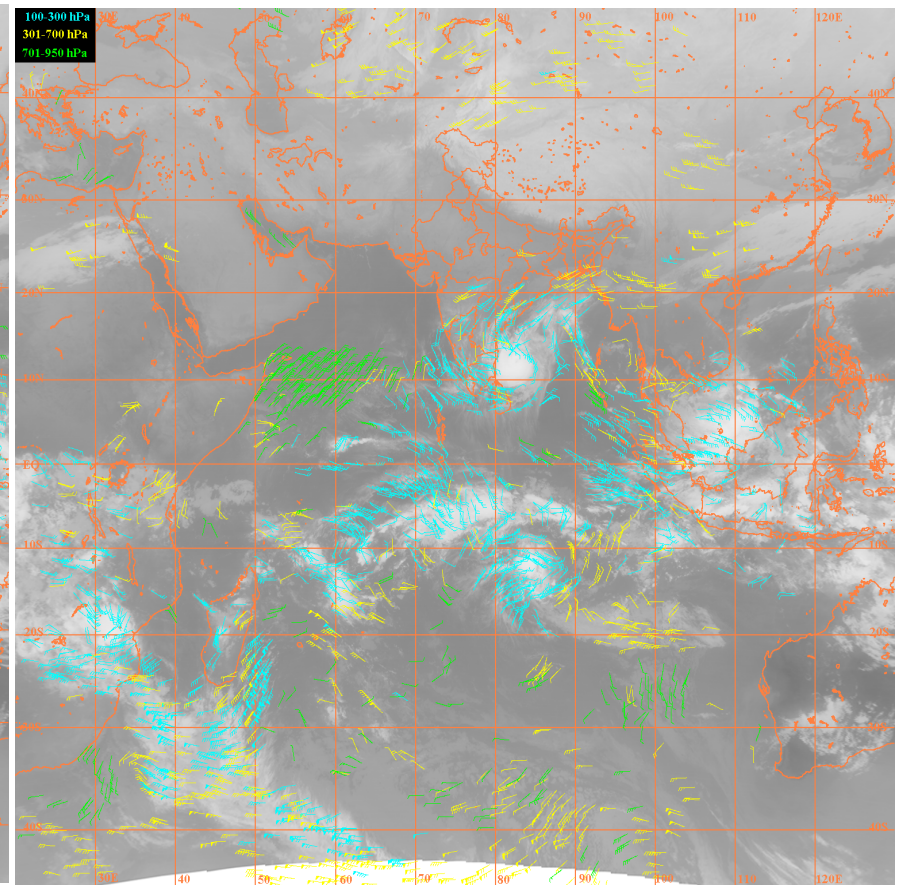


Cyclone Thane: IR winds - 29th December 2011: 00 UTC

Kalpana-1/ISRO

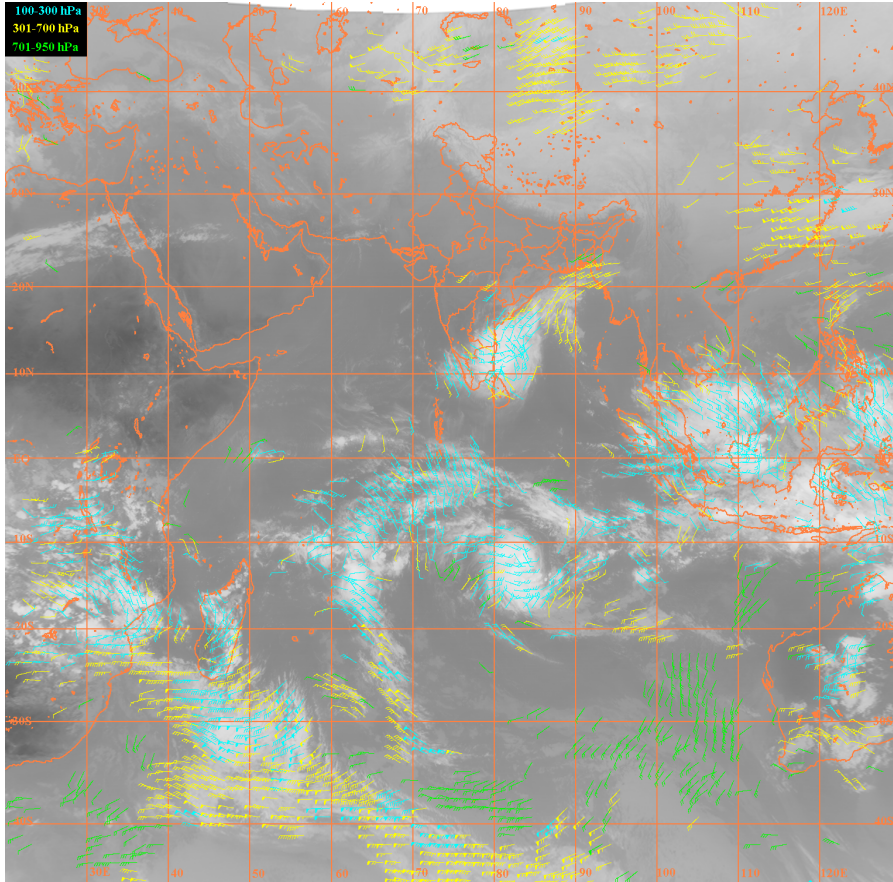


Meteosat7/CIMSS

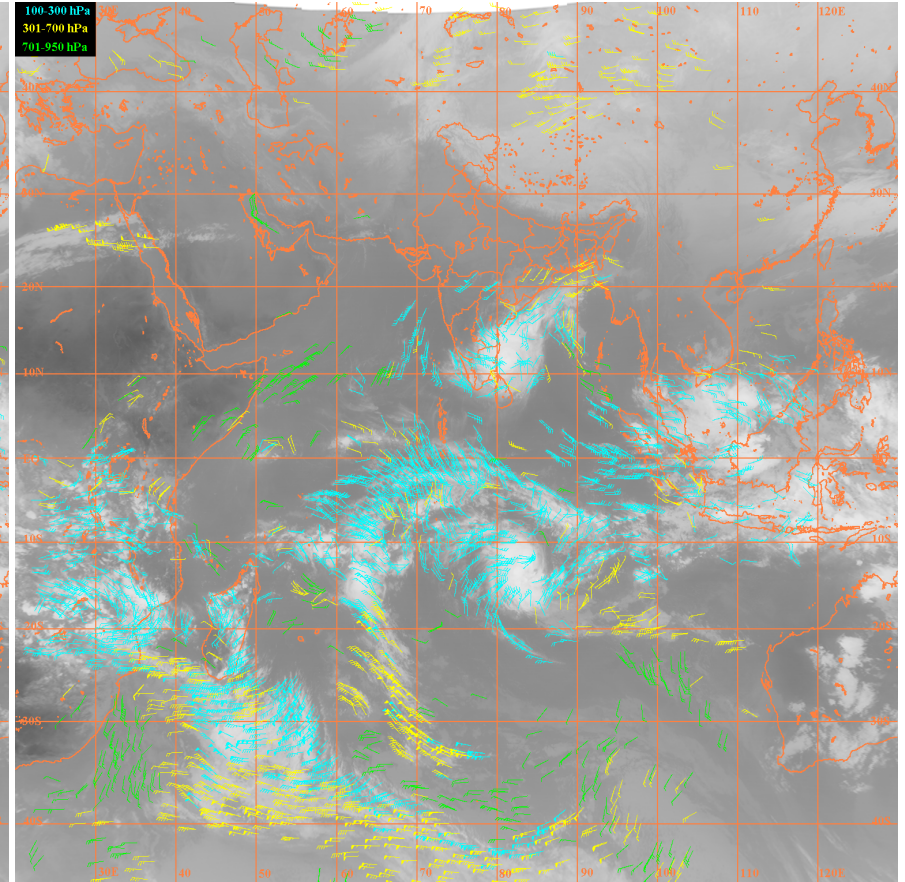


Cyclone Thane: IR winds - 29th December 2011: 12 UTC

Kalpana-1/ISRO

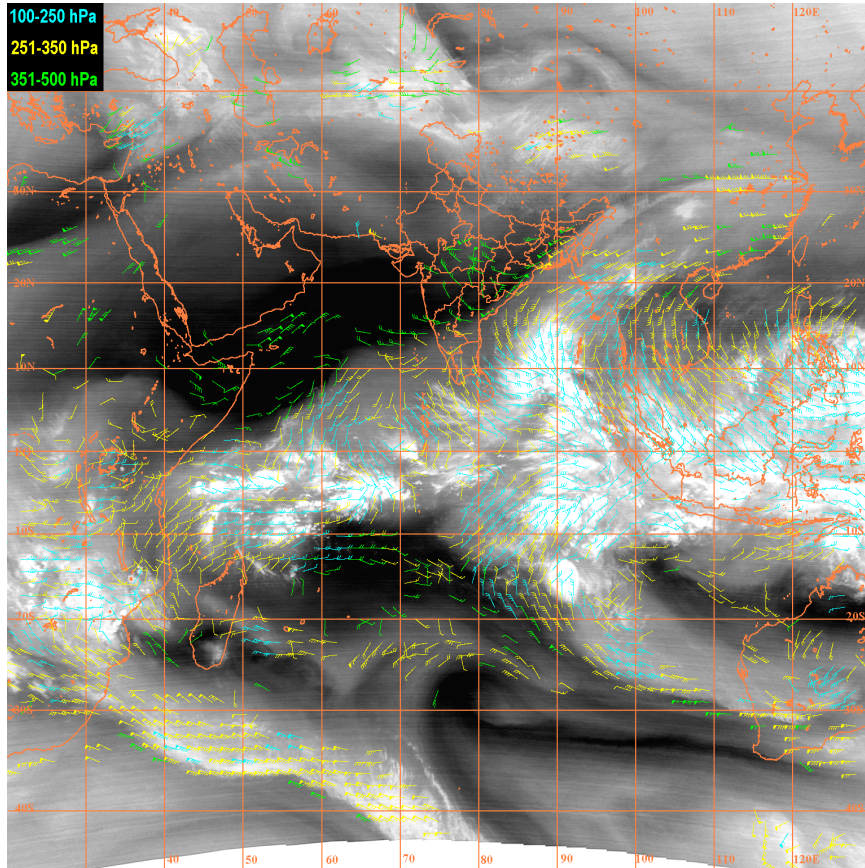


Meteosat7/CIMSS

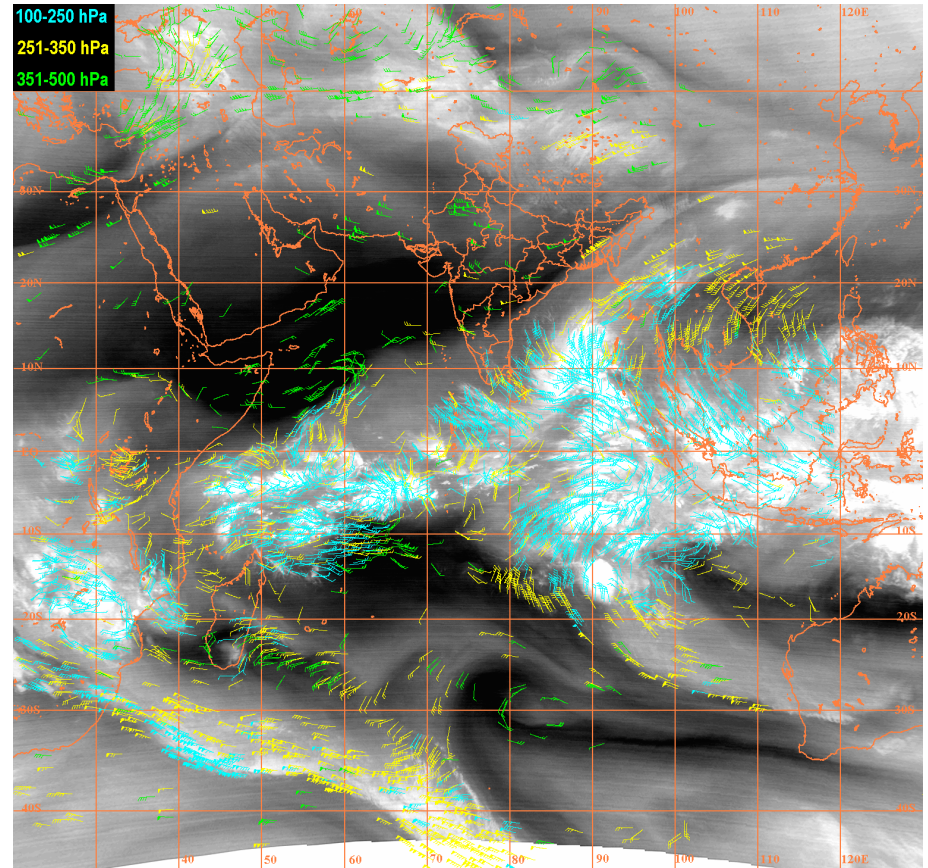


Cyclone Thane: WV winds - 25th December 2011: 00 UTC

Kalpana-1/ISRO

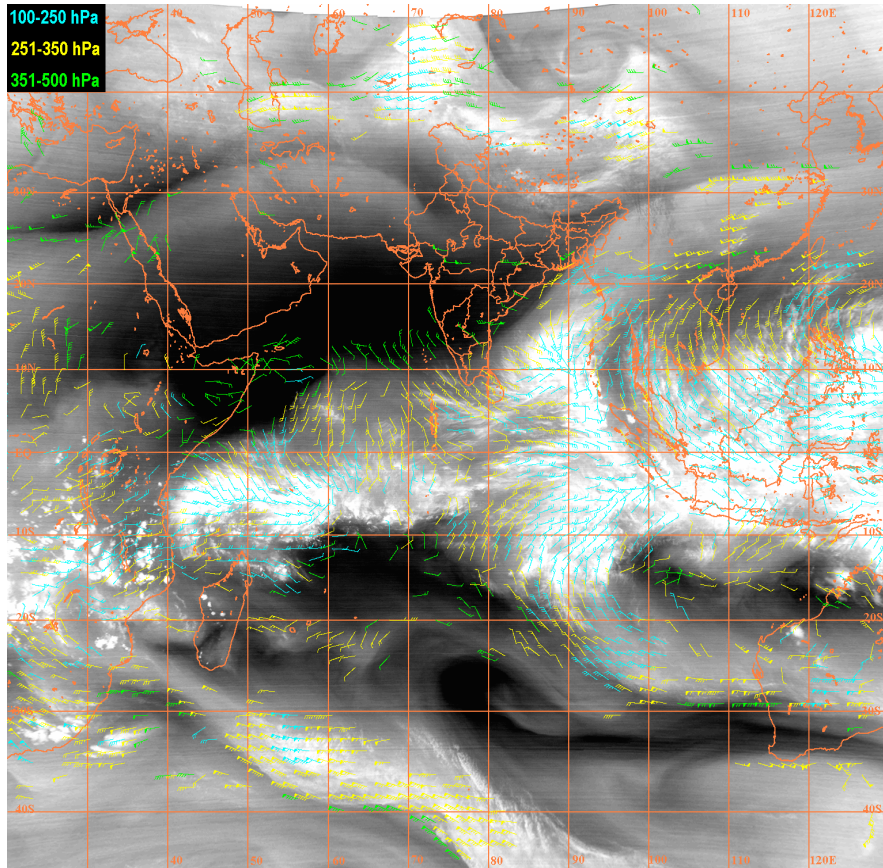


Meteosat7/CIMSS

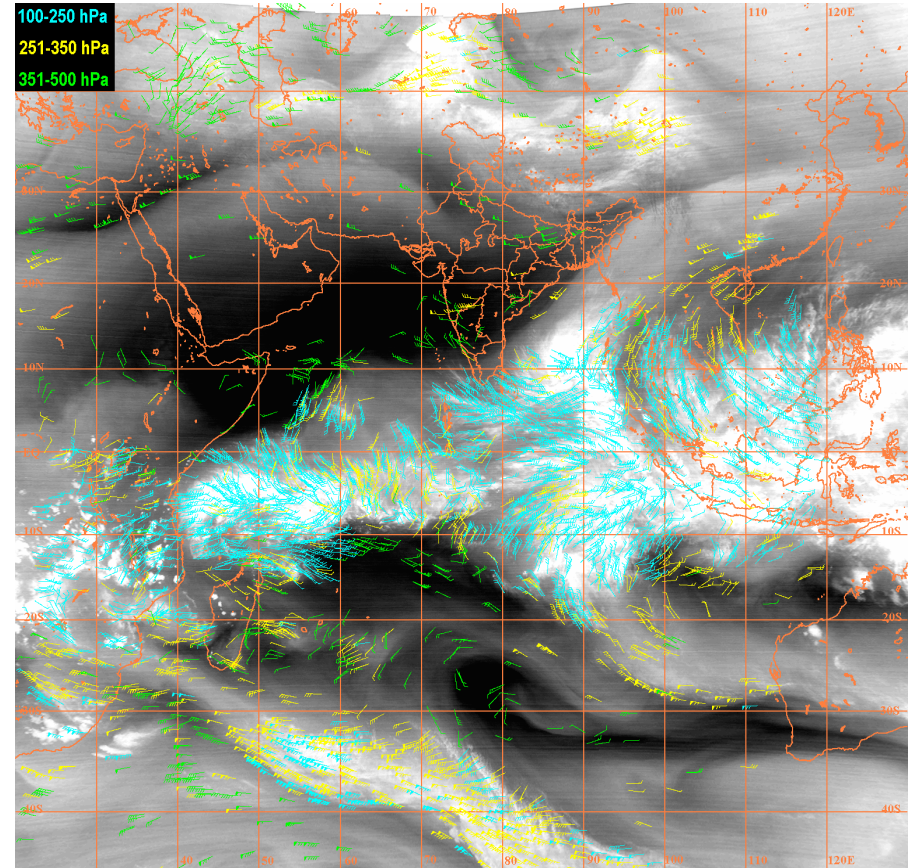


Cyclone Thane: WV winds - 25th December 2011: 12 UTC

Kalpana-1/ISRO

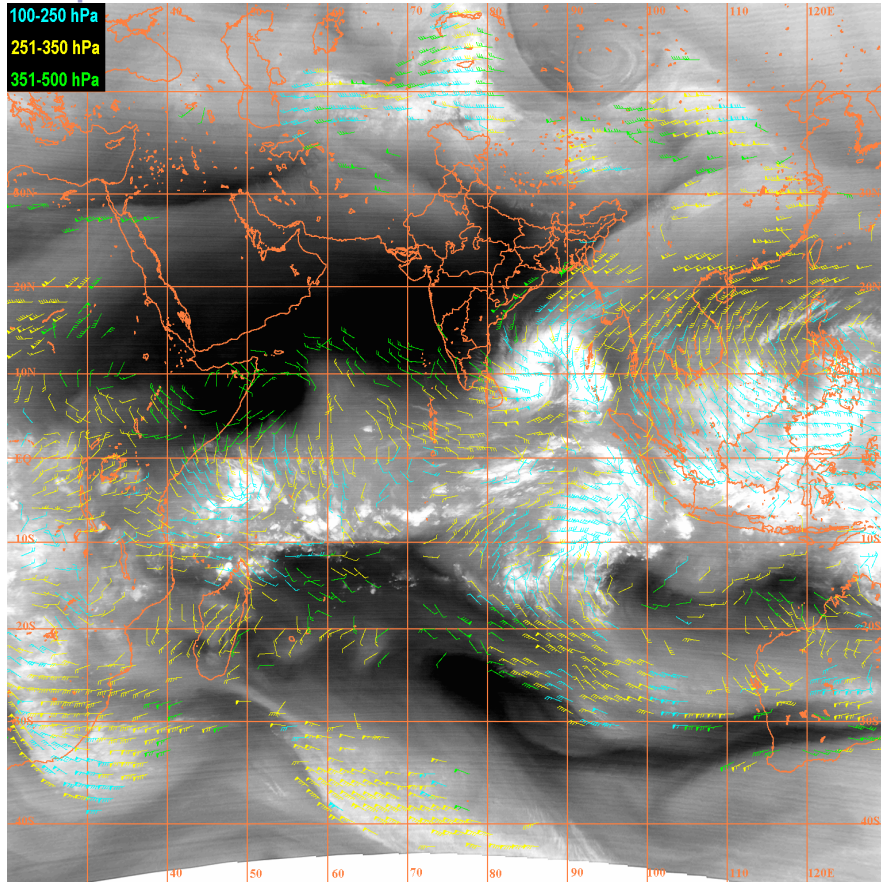


Meteosat7/CIMSS

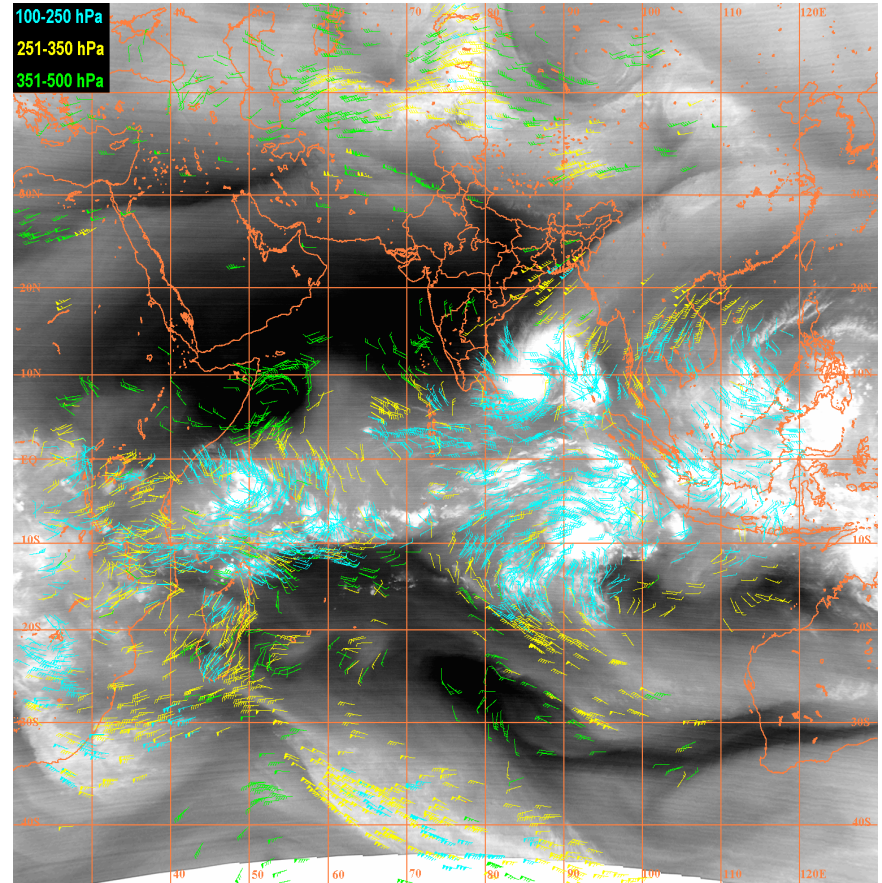


Cyclone Thane: WV winds - 26th December 2011: 00 UTC

Kalpana-1/ISRO

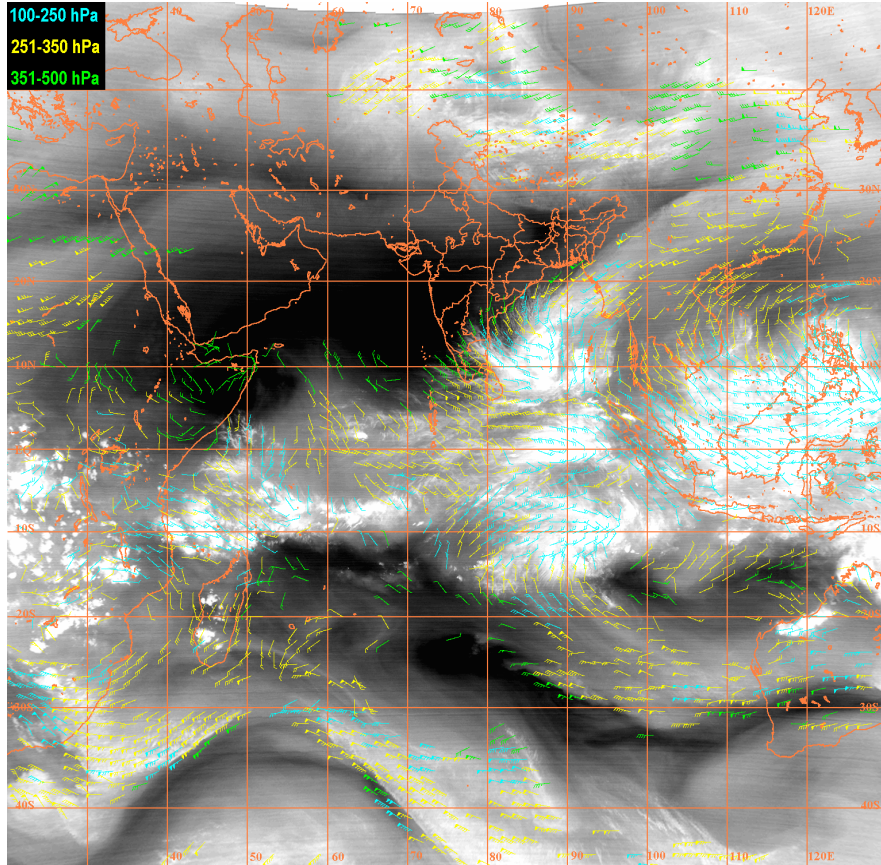


Meteosat7/CIMSS

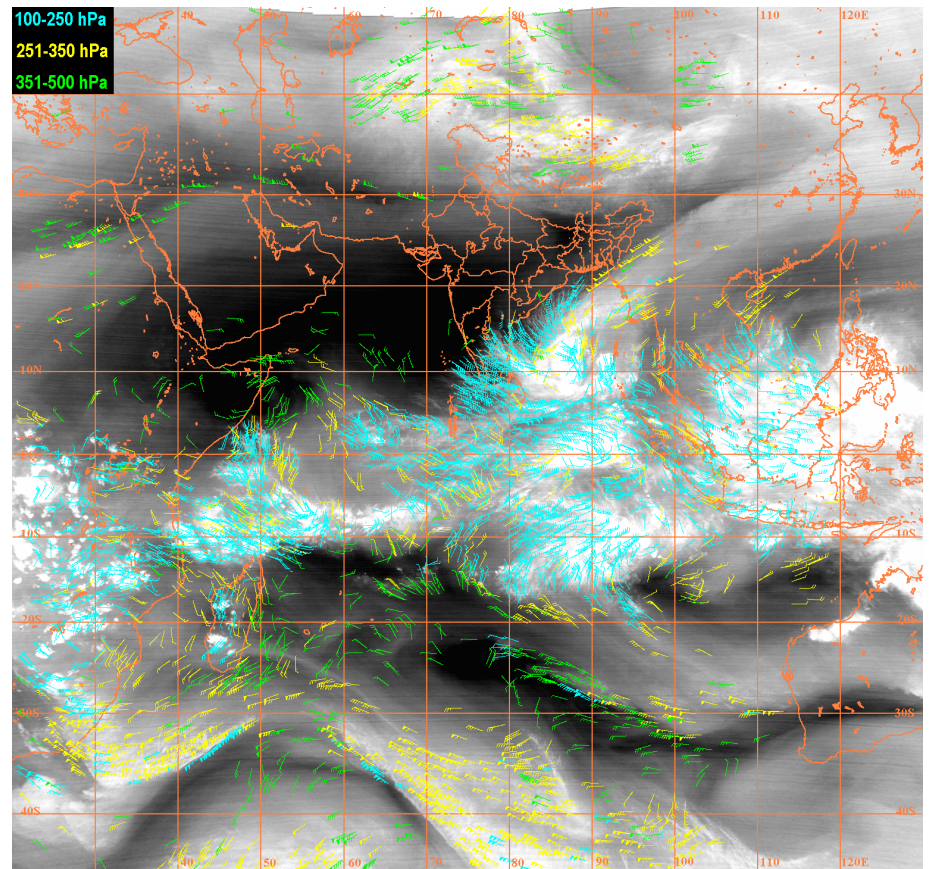


Cyclone Thane: WV winds - 26th December 2011: 12 UTC

Kalpna-1/ISRO

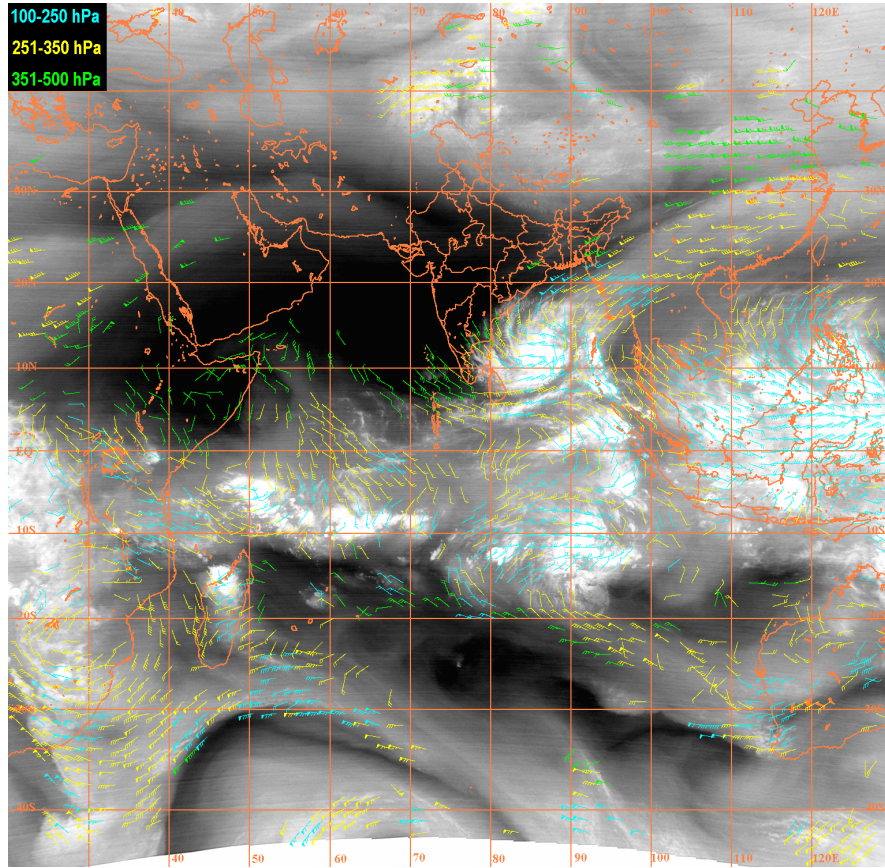


Meteosat7/CIMSS

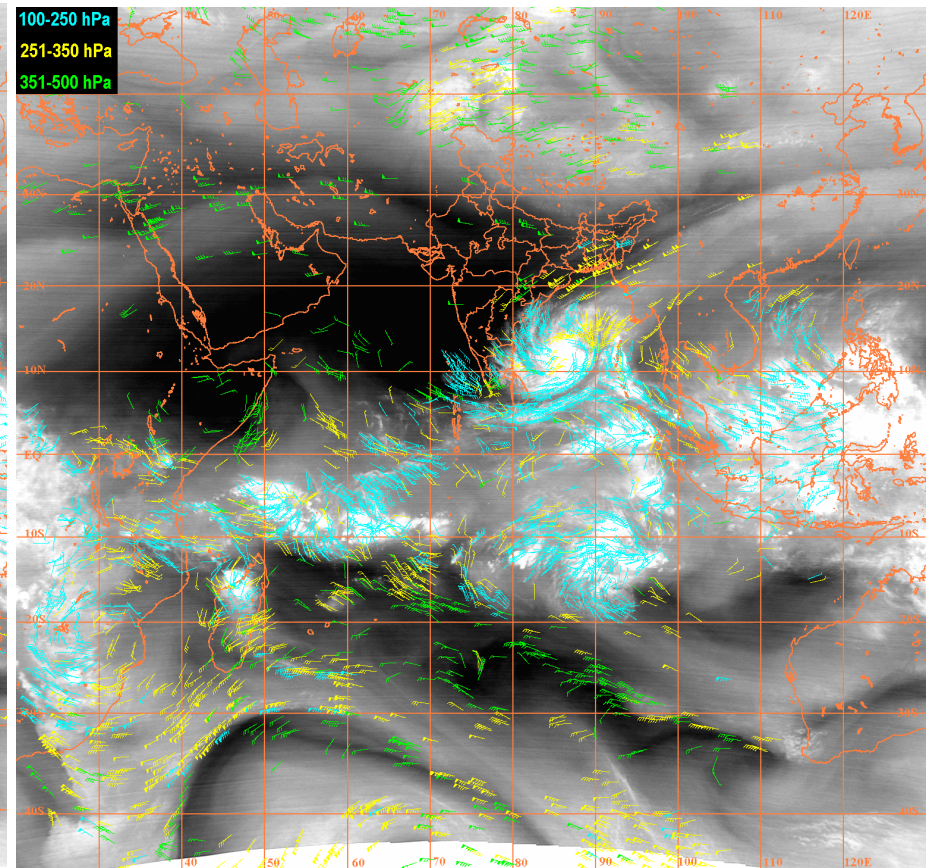


Cyclone Thane: WV winds - 27th December 2011: 00 UTC

Kalpana-1/ISRO

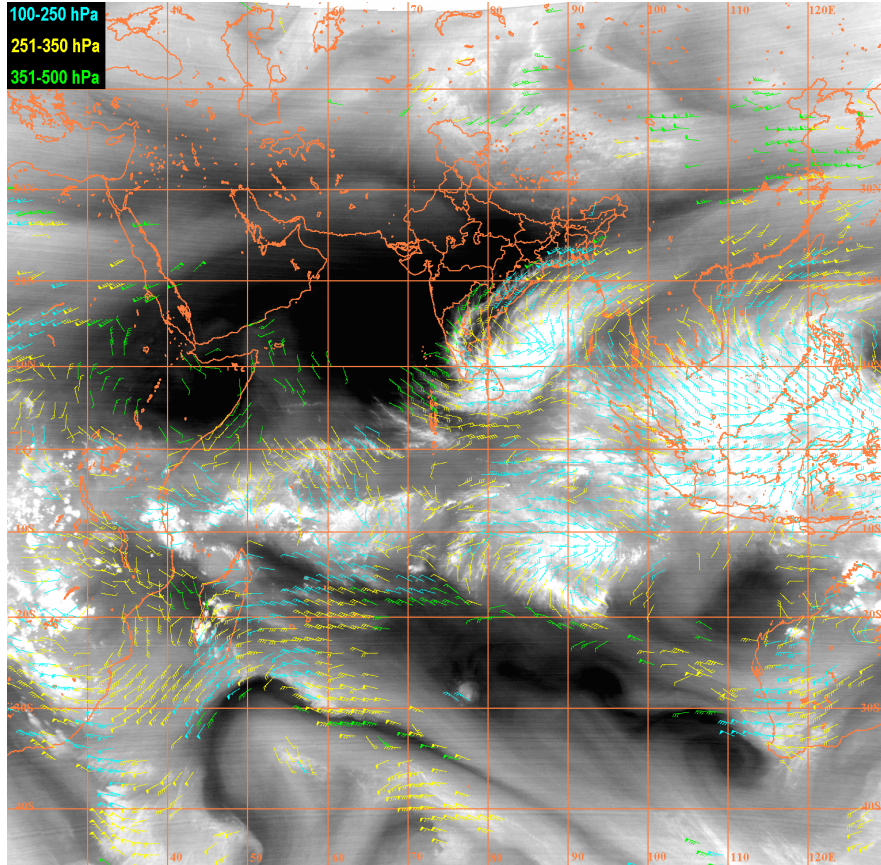


Meteosat7/CIMSS

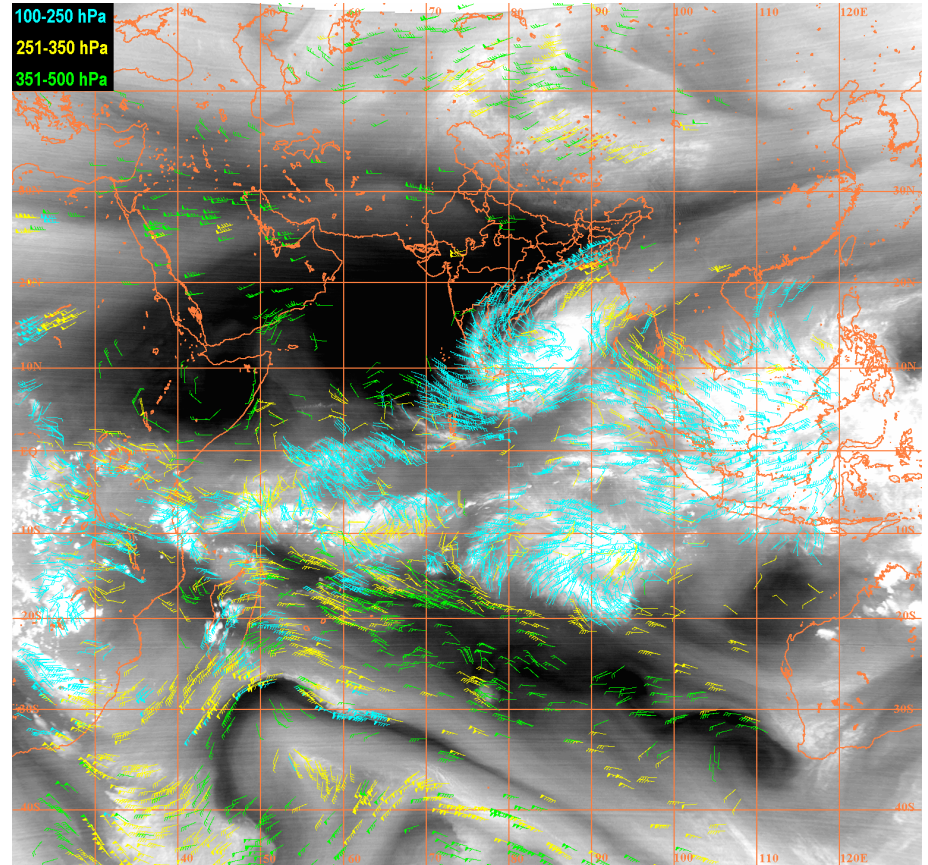


Cyclone Thane: WV winds - 27th December 2011: 12 UTC

Kalpana-1/ISRO

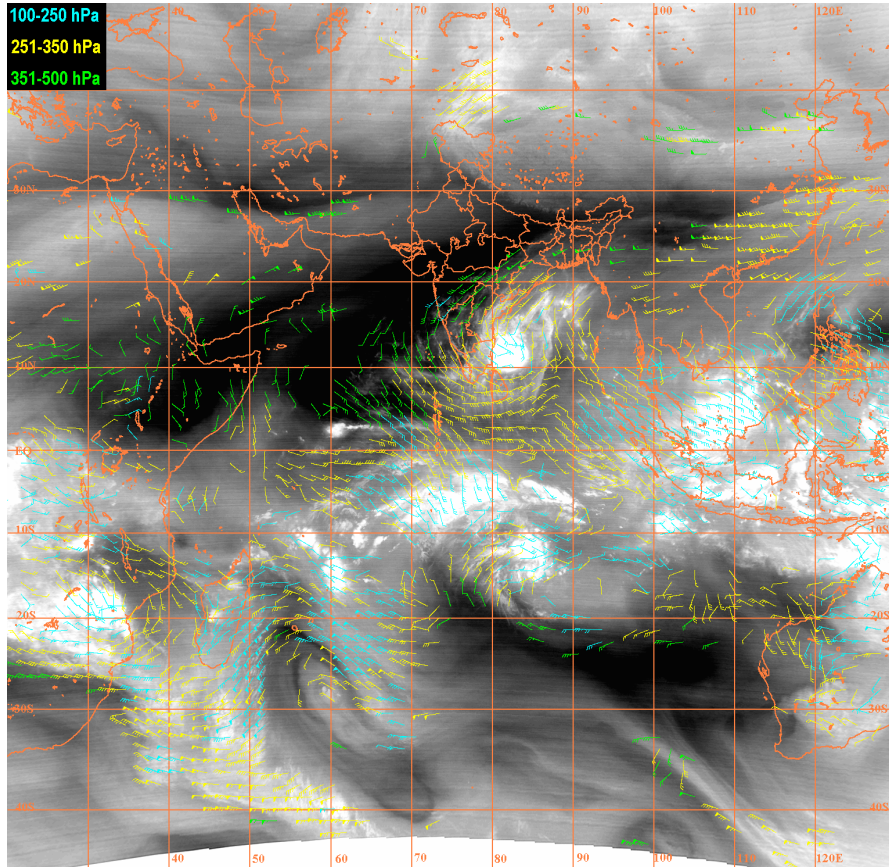


Meteosat7/CIMSS

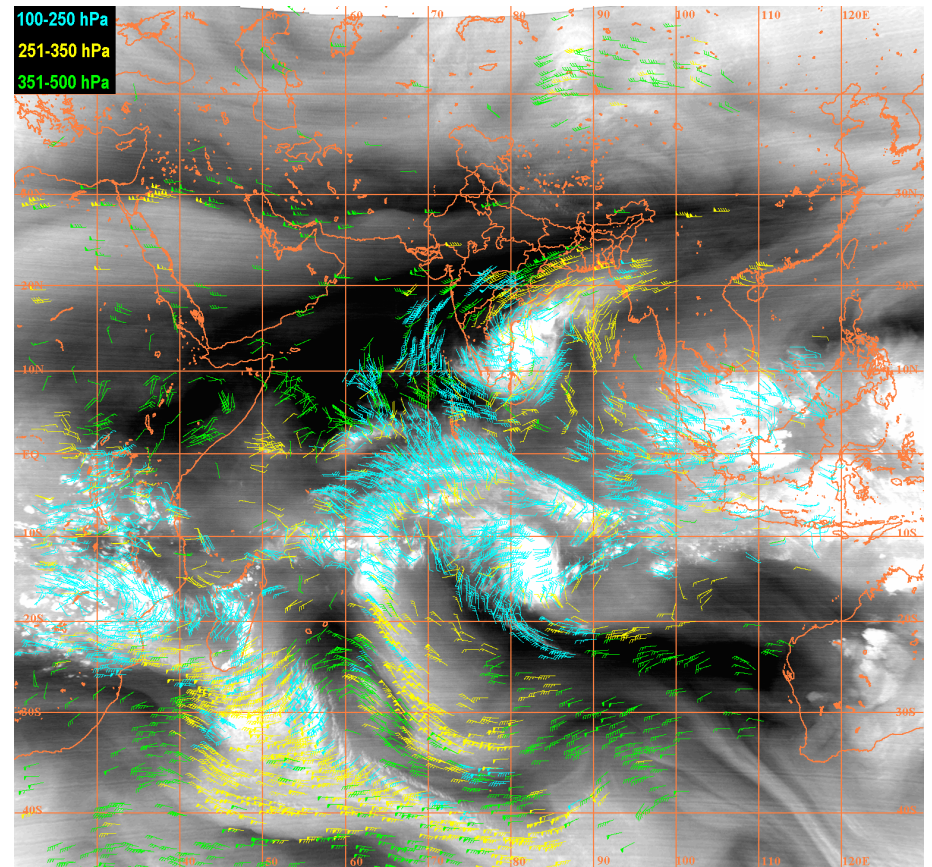


Cyclone Thane: WV winds - 29th December 2011: 00 UTC

Kalpana-1/ISRO

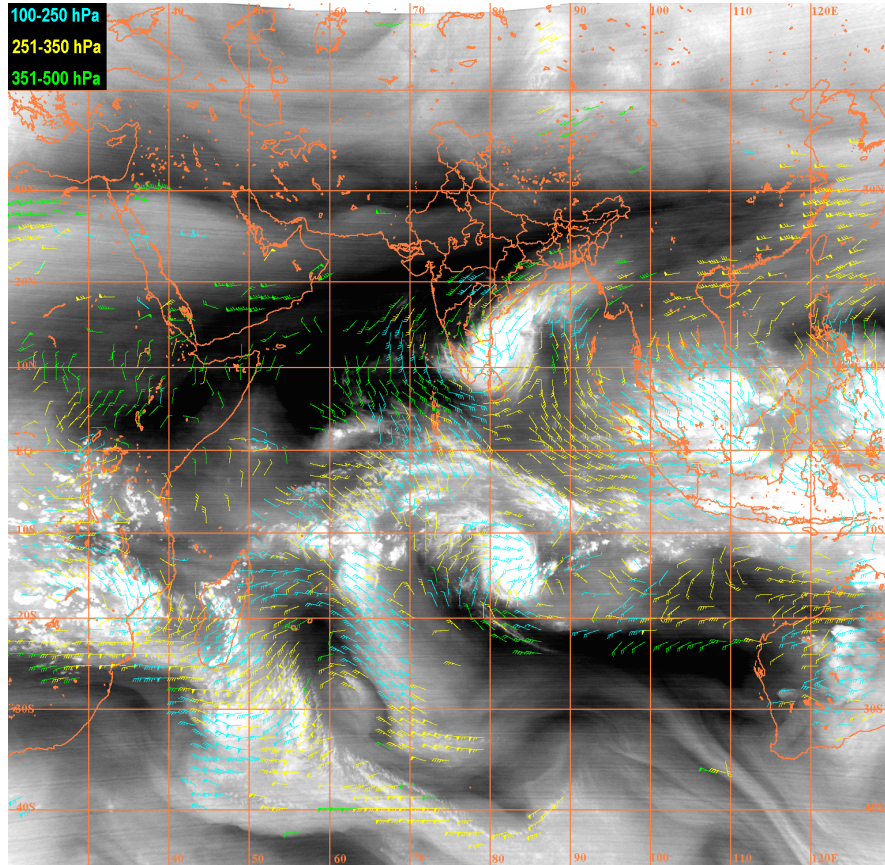


Meteosat7/CIMSS



Cyclone Thane: WV winds - 29th December 2011: 12 UTC

Kalpana-1/ISRO



Meteosat7/CIMSS

